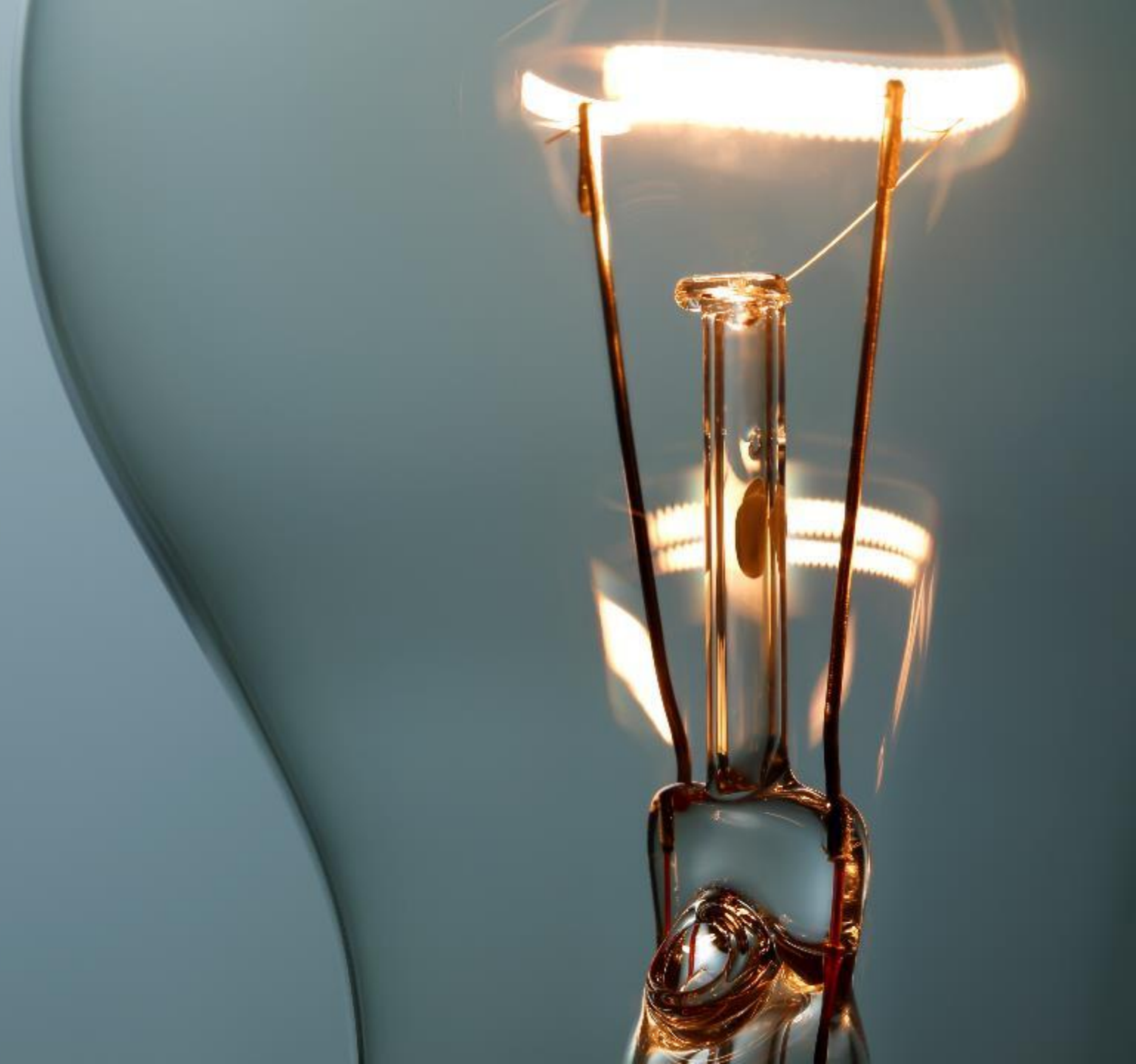


Compton effect

Ghofrane Bel-Hadj-Aissa



Contents

I- History of the quantization of matter

II- The discovery of the electron

III- Scattering of electromagnetic radiation by electrons

IV- Photon-electron scattering

V- Conclusions

I. History of the quantization of matter

5th century B.C



Democritus

17th century



Pierre Gassendi



Robert Hooke



Isaac Newton

18th century



John Dalton

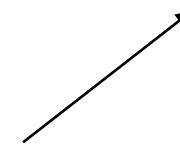
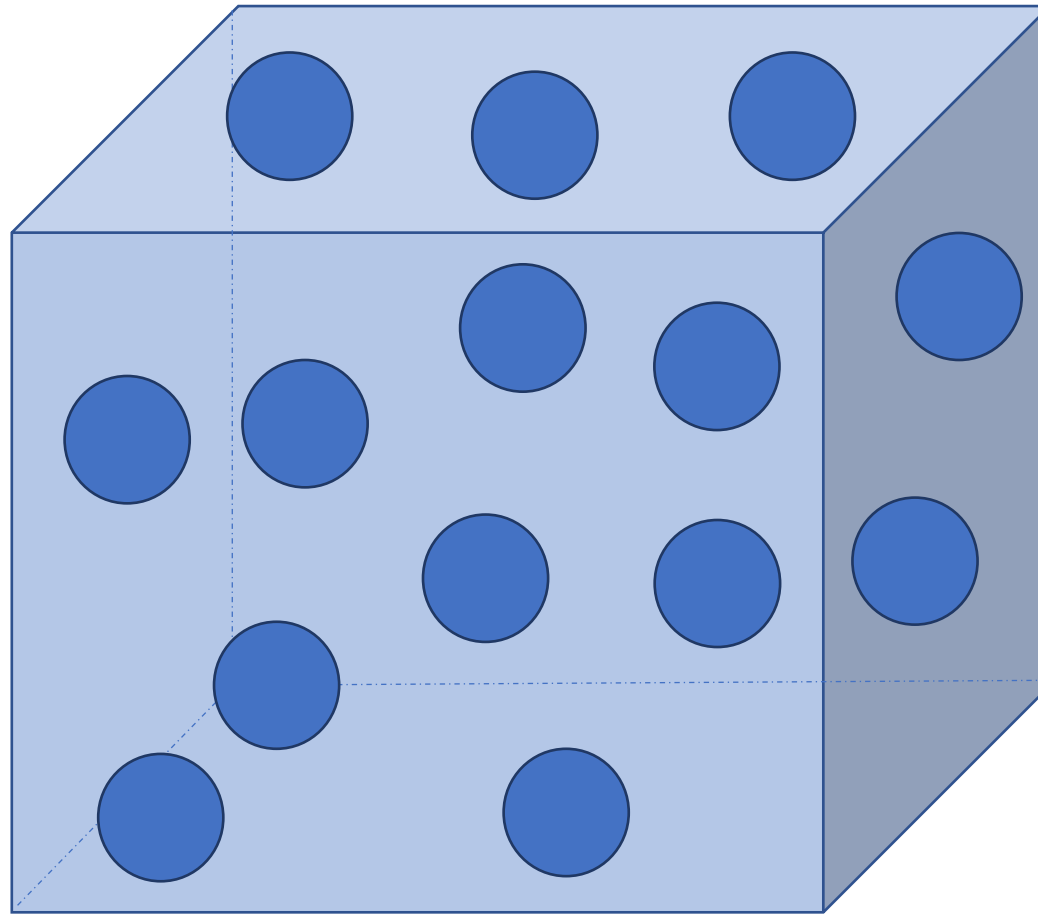


Amedeo Avogadro



Antoine Lavoisier

Dalton's atomic model



All matter consists of
Tiny indivisible particles

II- The discovery of the electron

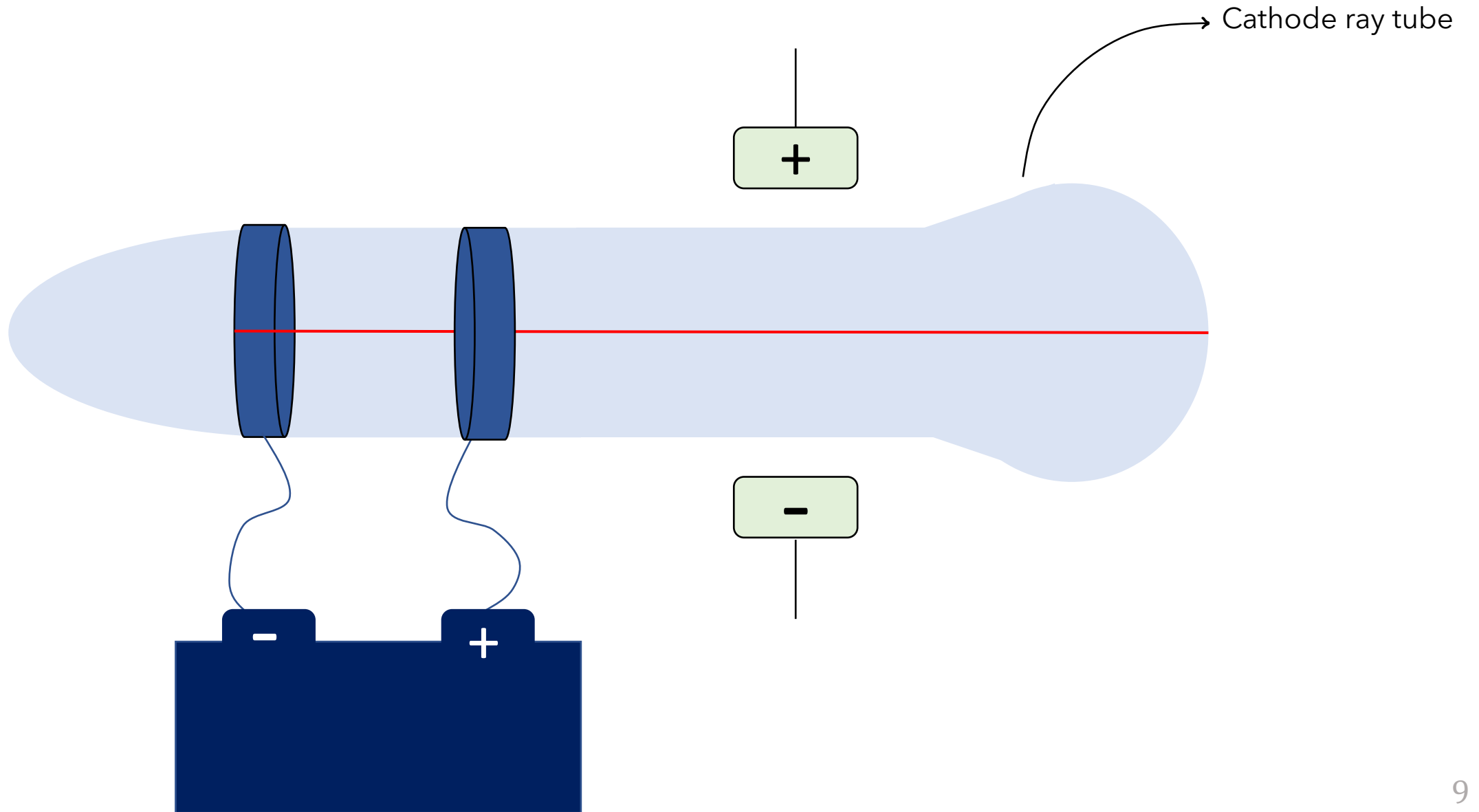
Joseph John Thomson's experiment

Joseph John Thomson
1897

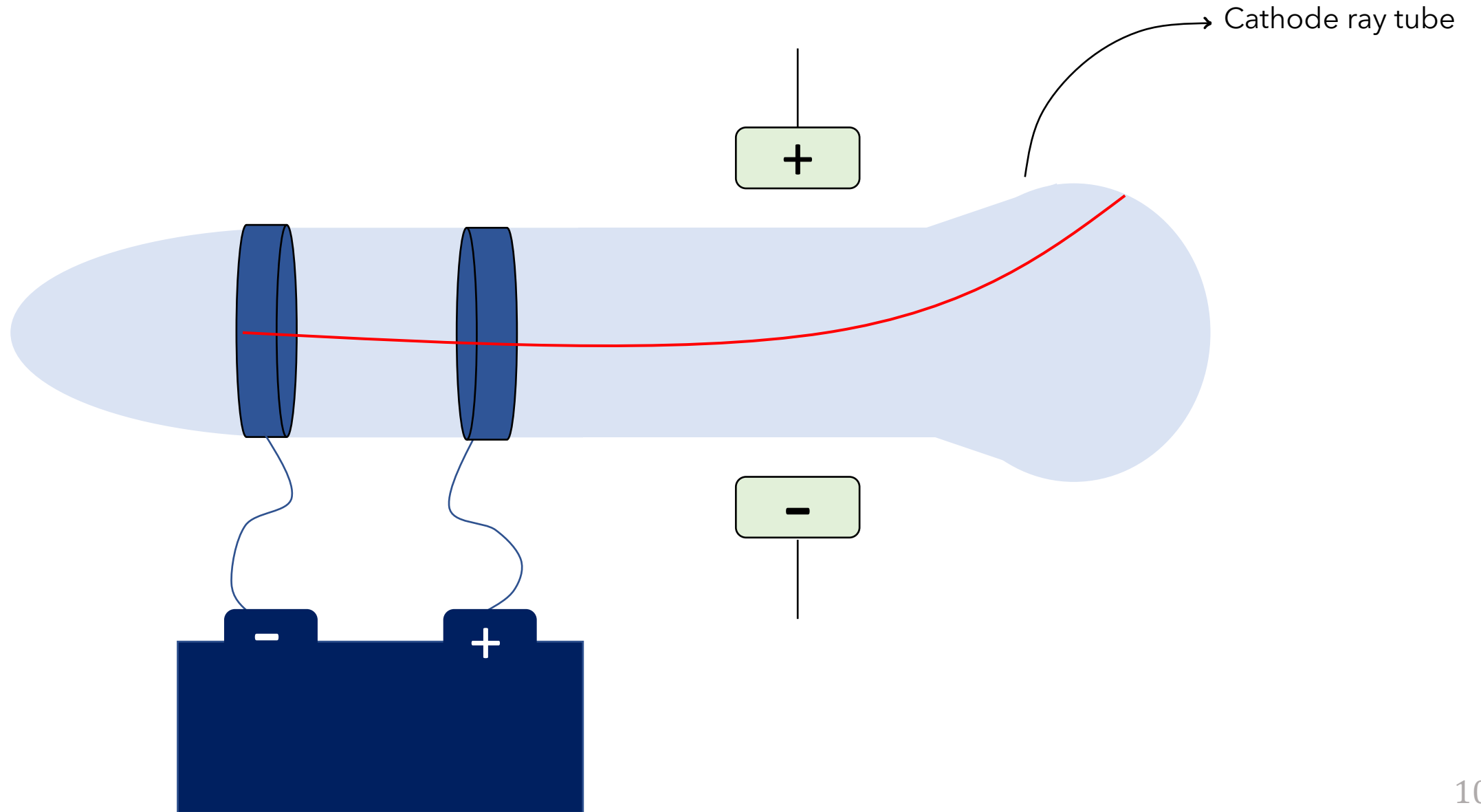
The cathode ray tube experiment:
Discovery of the electron



Experimental set-up of Thomson's experiment



Experimental set-up of Thomson's experiment



- Cathode rays must be made up of substance that is negatively charged
- Particles that make up cathode rays are 1000 times smaller than a Hydrogen atom
- All different metals give off cathode ray



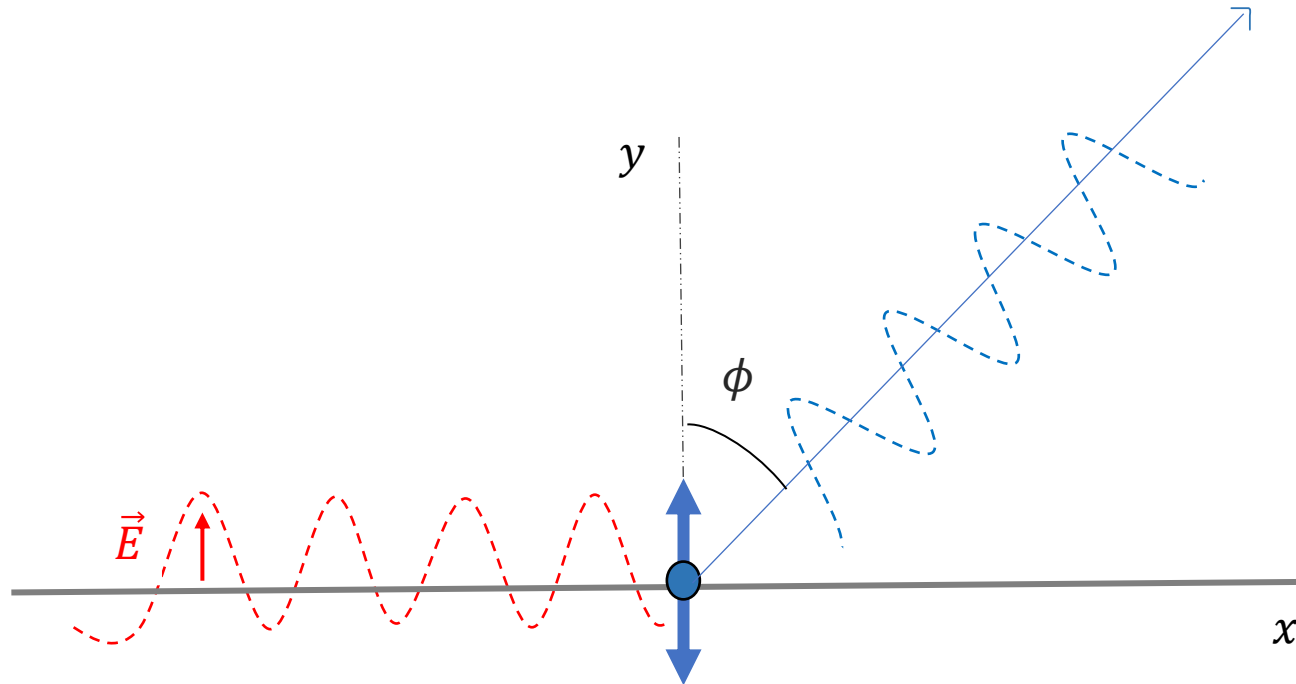
Atoms have tiny, negatively charged particles inside them:

Electrons

III- Scattering of electromagnetic radiation by electrons

Thomson's scattering **1904**

Thomson scattering is the elastic scattering of low energy electromagnetic radiation by a free charged particle



$$\vec{E}(\mathbf{r}, t) = e \left(\frac{\vec{n} - \vec{\beta}}{\gamma^2 (1 - \vec{\beta} \cdot \vec{n})^3 r^2} \right) + \frac{e}{c} \left(\frac{\vec{n} \wedge [(\vec{n} - \vec{\beta}) \wedge \dot{\vec{\beta}}]}{\gamma^2 (1 - \vec{\beta} \cdot \vec{n})^3 r} \right)$$



$$\sim \frac{1}{r^2} \ll 1$$



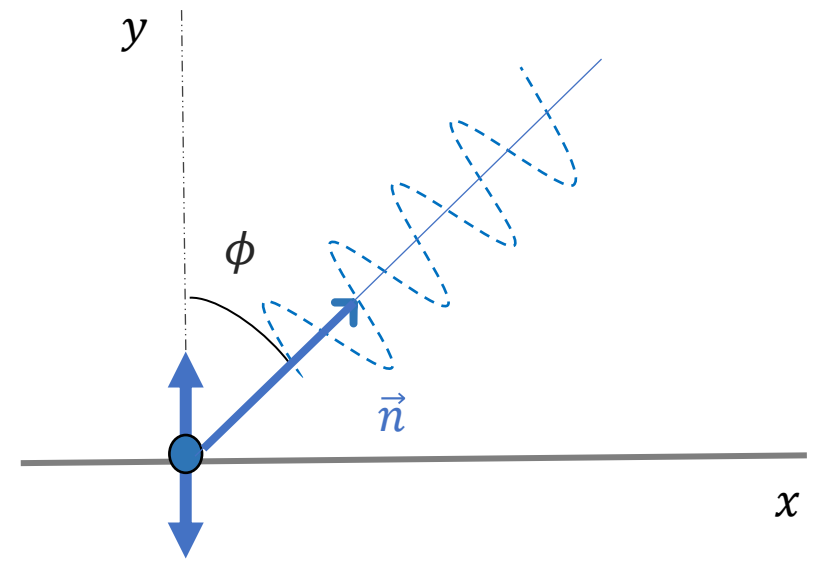
$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} \quad \dot{\vec{\beta}} = \frac{\vec{a}}{c}$$

$$\vec{\beta} = \frac{\vec{v}}{c} \ll 1$$

$$\vec{B}(\mathbf{r}, t) = \frac{\vec{n}}{c} \wedge \vec{E}(\mathbf{r}, t)$$

$$\vec{E}(\mathbf{r}, t) = \frac{e}{c} \frac{\vec{n} \wedge (\vec{n} \wedge \dot{\vec{\beta}})}{r}$$

$$\vec{B}(\mathbf{r}, t) = \vec{n} \wedge \vec{E} = \frac{e}{c^2} \frac{(\vec{n} \wedge \dot{\vec{\beta}})}{r}$$



- Equation of motion

$$m_e \vec{a} = e \vec{E} + e \underbrace{\vec{v} \wedge \vec{B}}_{\ll 1}$$

$$\Rightarrow \vec{a} = \frac{e}{m_e} \vec{E}$$

$$\vec{E} = E_0 \sin(\omega_0 t) \vec{y}$$

$$\Rightarrow a_y = \frac{e}{m_e} E_0 \sin(\omega_0 t) \quad \text{Acceleration}$$

$$\Rightarrow y = \iint \frac{eE_0}{m_e} \sin(\omega_0 t) = \frac{eE_0}{\omega_0^2 m_e} \sin(\omega_0 t)$$

$$\Rightarrow \vec{d} = e \vec{y} = \frac{e^2 E_0}{\omega_0^2 m_e} \sin(\omega_0 t) \quad \text{Dipole moment}$$

- Cross section

$$\sigma = \frac{R}{I}$$

σ = Cross section

R = Number of reactions per unit time
per nucleus

I = Number of incident particles per unit time
per unit area

$$\sigma = \frac{P}{\langle S \rangle}$$

P = Power radiate by the charged particle

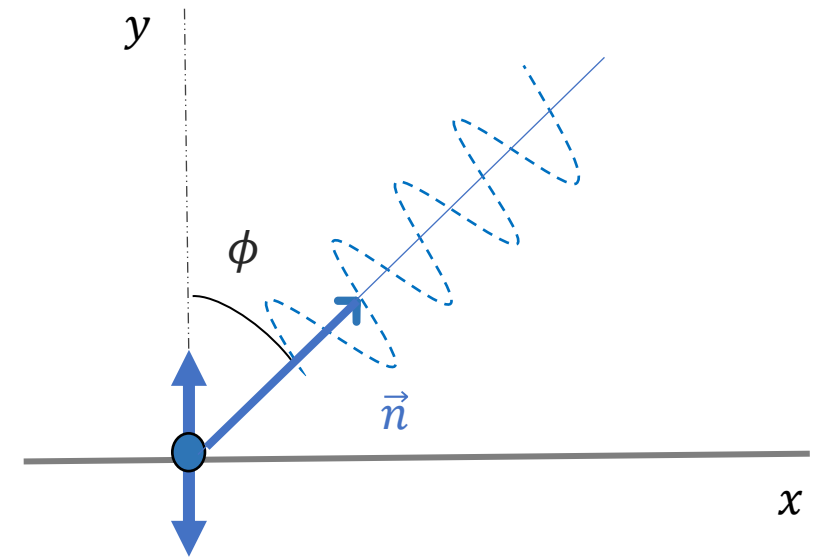
$\langle S \rangle$ = The incident flux (the time-average
Poynting vector)

Differential cross section for Thomson's scattering

$$\frac{d\sigma}{d\Omega} = \frac{dP}{d\Omega} \frac{1}{\langle S \rangle}$$

$$\Rightarrow \frac{dP}{d\Omega} = \frac{dP}{dA_n} r^2 = r^2 (\vec{n} \cdot \vec{S})$$

$$\Rightarrow \frac{dP}{d\Omega} = \frac{e^4 E_0^2}{4m_0^2 c^3} \sin^2 \phi$$



$$\frac{d\sigma}{d\Omega} = \frac{e^4}{m_0^2 c^4} \sin^2 \phi$$



Thomson Cross section (independent from the frequency)

IV- Photon-electron scattering

Compton scattering

- Compton scattering, is the decrease in energy (increase in wavelength) of an X-ray or gamma ray photon, when it interacts with matter. It is the high frequency limit of Thomson scattering.
- Thomson scattering, the classical theory of charged particles scattered by an electromagnetic wave, cannot explain any shift in wavelength.

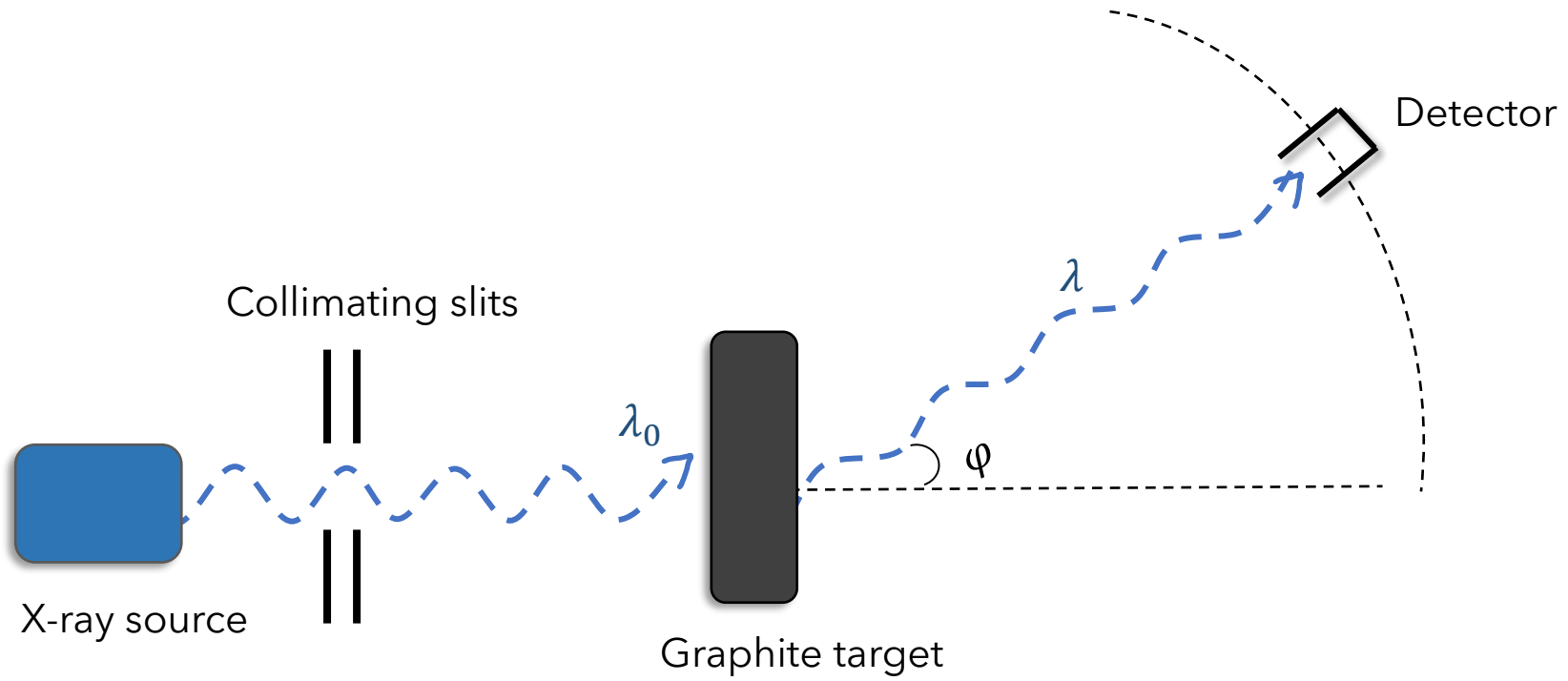


Compton embarked upon a more radical explanation based on the (at the time) new quantum theory.

Arthur Compton
1923

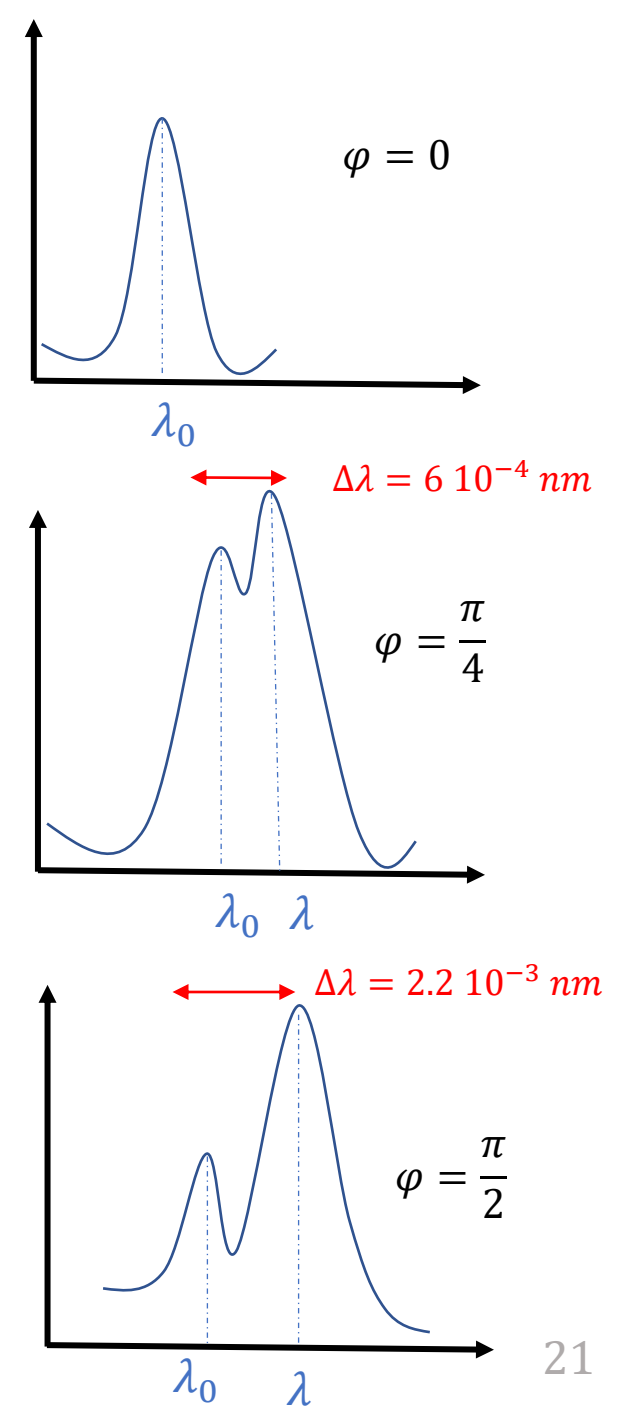


Experimental set-up of Compton scattering **1923**



Compton found that:

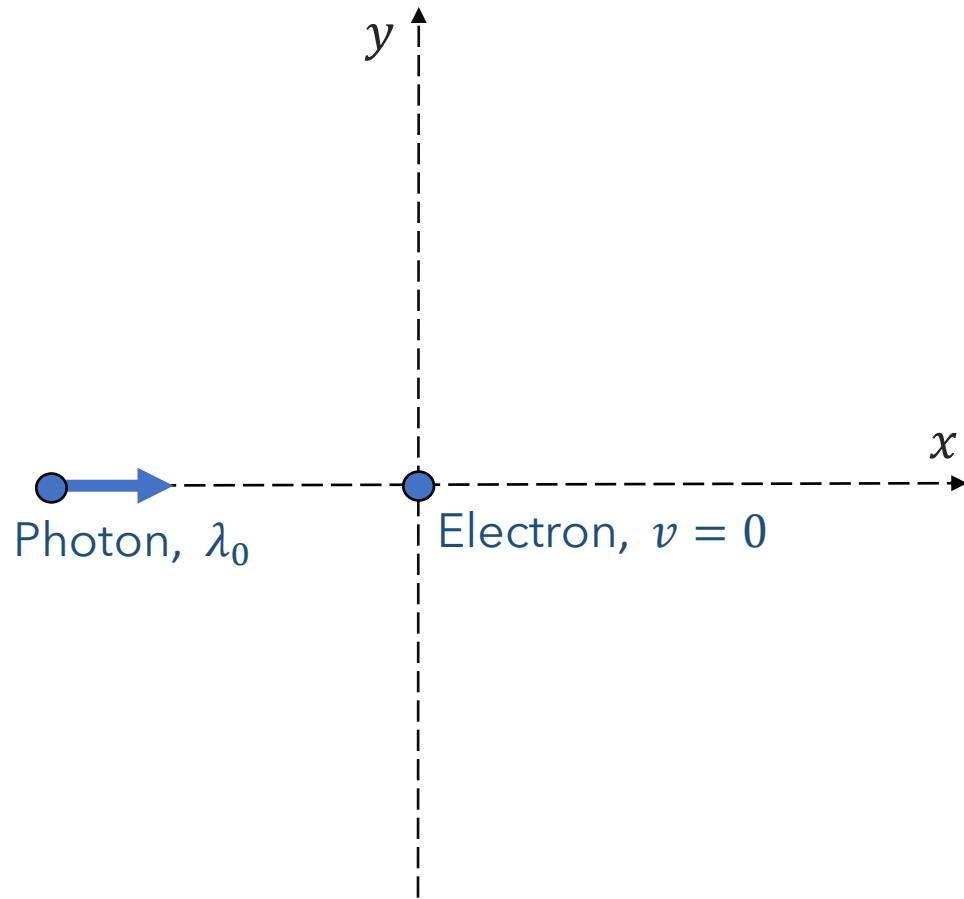
- One component had a wavelength λ_0 equal to that of the incident radiation and second component had a wavelength $\lambda = \lambda_0 + \Delta\lambda$.
- $\Delta\lambda$ varies with the scattering angle
- $\Delta\lambda$ increases rapidly at large scattering angles
- $\Delta\lambda$ is independent of the incident wavelength λ_0
- $\Delta\lambda$ is independent of the scattering material



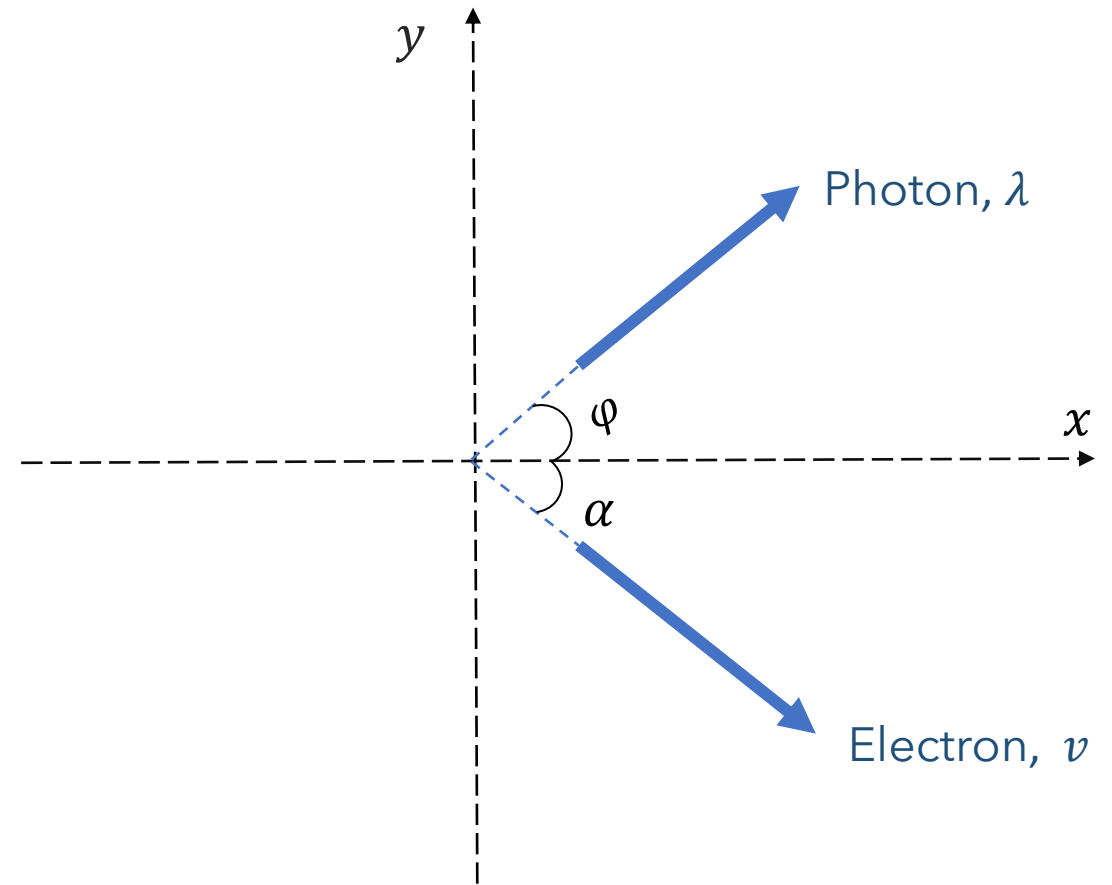
Compton explained and modeled the data by :

- 1)** Assuming that the electron is free or loosely bound to the valence shell of atoms.
- 2)** Assuming a particle (photon) nature for light.
- 3)** Applying conservation of energy and conservation of momentum to the collision between the photon and the electron.

Derivation of the Compton scattering calculations



Pre-collision



Post-collision

Conservation of energy

$$\begin{aligned} h\nu_0 + m_0c^2 &= h\nu + mc^2 \\ \textcircled{1} \quad \frac{h}{\lambda_0} - \frac{h}{\lambda} + m_0c &= mc \end{aligned} \quad \begin{array}{l} \curvearrowright \\ c = \nu\lambda \end{array}$$

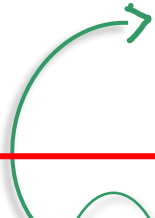
Conservation of momentum

$$\begin{aligned} \textcircled{2} \quad x\text{-component} \quad \frac{h}{\lambda_0} &= \frac{h}{\lambda} \cos\varphi + mv \cos\alpha \\ y\text{-component} \quad 0 &= \frac{h}{\lambda} \sin\varphi - mv \sin\alpha \end{aligned}$$

$$\textcircled{1} \Rightarrow \frac{h^2}{\lambda_0^2} + \frac{h^2}{\lambda^2} - \frac{2h^2}{\lambda_0\lambda} - 2m_0c \left(\frac{h}{\lambda_0} - \frac{h}{\lambda} \right) + m_0^2c^2 = m^2c^2 \quad \textcircled{3}$$

$$\textcircled{2} \Rightarrow \frac{h^2}{\lambda_0^2} + \frac{h^2}{\lambda^2} - \frac{2h^2}{\lambda\lambda_0} \cos\varphi = m^2v^2 \quad \textcircled{4}$$

Compton wavelength



$$\textcircled{3} - \textcircled{4} \Rightarrow \Delta\lambda = \lambda - \lambda_0 = \frac{h}{m_0c} (1 - \cos\varphi)$$

Compton shift

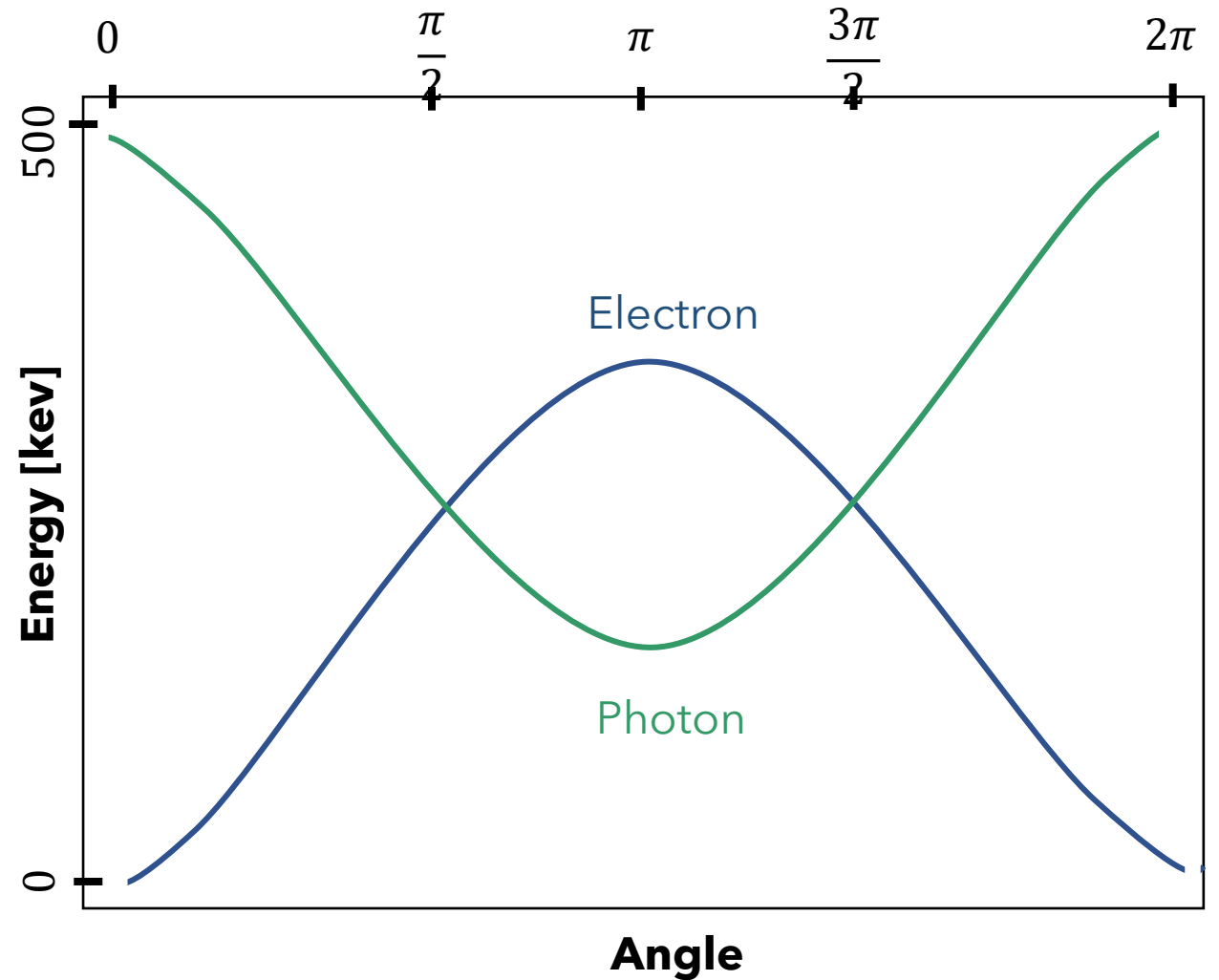
$$\Delta\lambda = \lambda - \lambda_0 = \frac{h}{m_0c} (1 - \cos \varphi)$$

$$\lambda = \frac{hc}{E}$$

$$\Rightarrow E_f = \frac{E_i}{1 + \frac{E_i (1 - \cos \varphi)}{m_e c^2}}$$

Final energy of the scattered photon is a function of:

- Initial energy of the photon
- Scattering angle



Energies of a photon at 500 keV and an electron after Compton scattering.

Klein-Nishina cross section

$$\frac{d\sigma}{d\Omega} = \frac{1}{2} \left(\frac{e^4}{m_0^2 c^4} \right)^2 \left(\frac{\lambda_0}{\lambda} \right)^2 \left(\left(\frac{\lambda_0}{\lambda} \right) + \left(\frac{\lambda}{\lambda_0} \right) - 2 \sin^2 \theta \cos^2 \phi \right)$$

Low energy limit $\lambda_0 \approx \lambda$

$$\frac{d\sigma}{d\Omega} = \left(\frac{e^4}{m_0^2 c^4} \right)^2 \sin^2 \phi$$

Thomson differential cross section

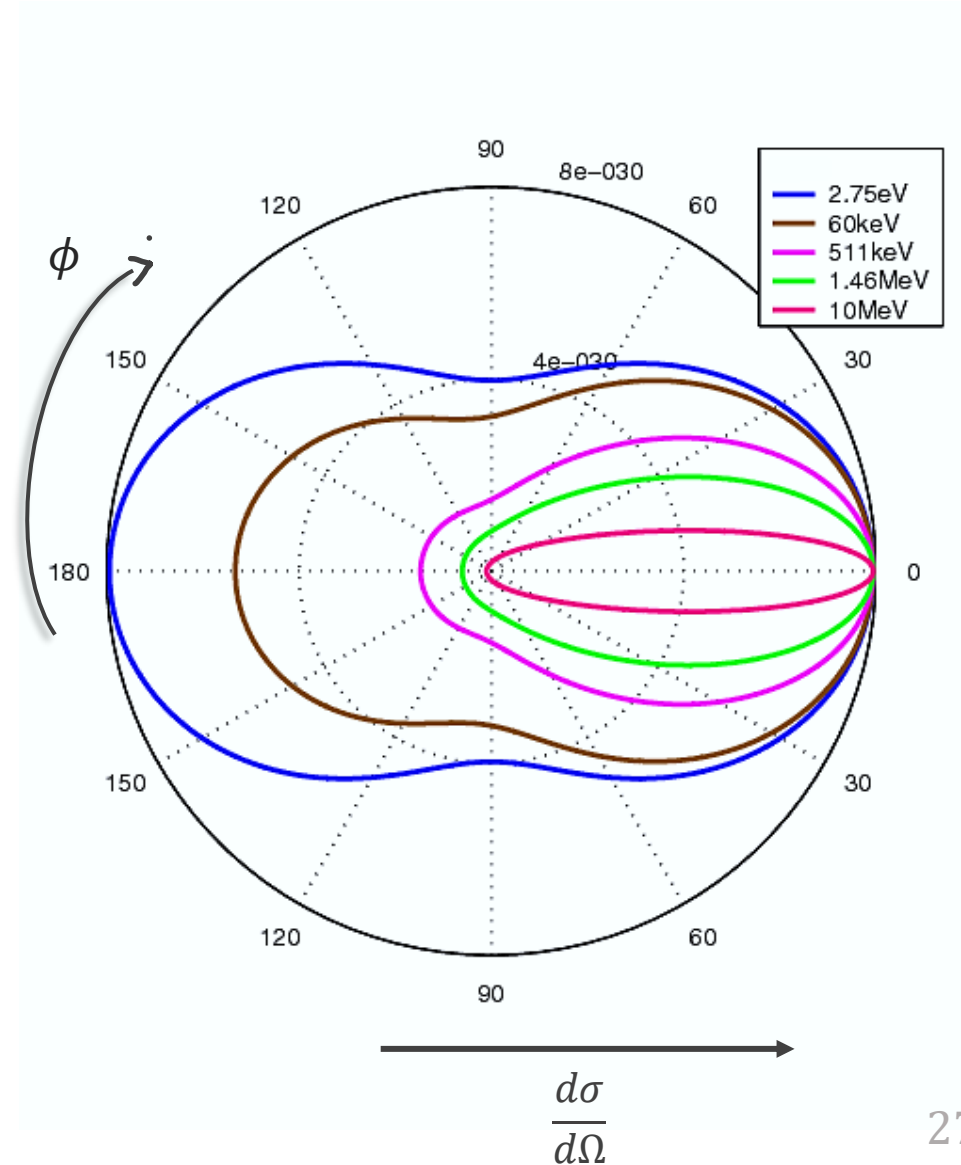


Oskar Klein

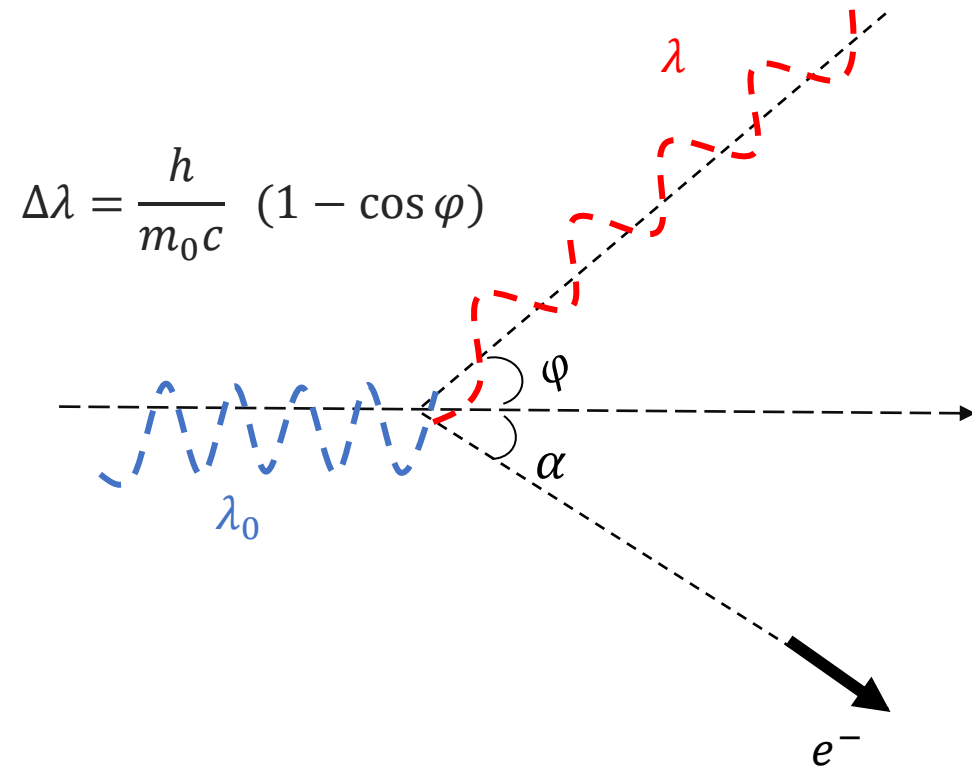


Yoshio Nishina

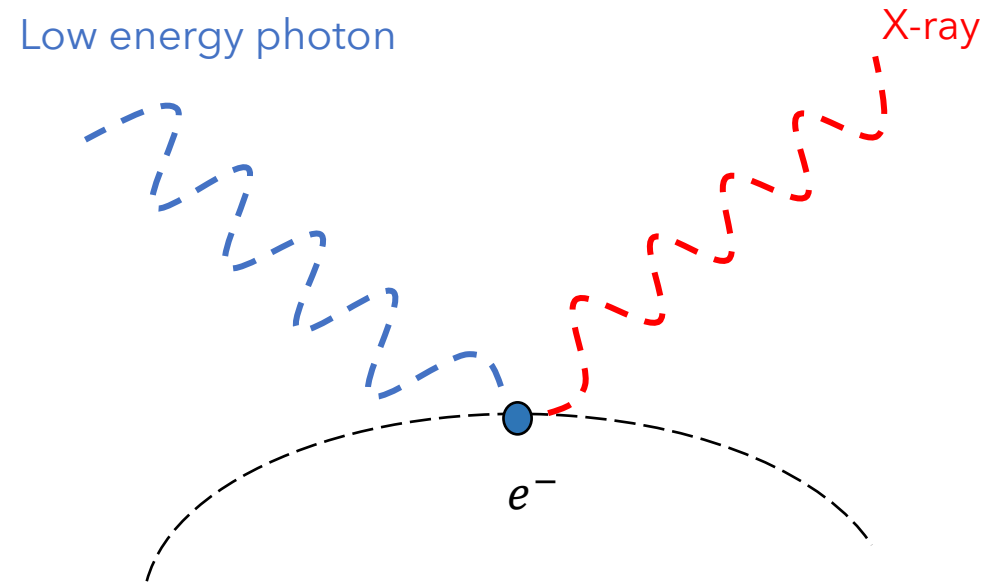
1929



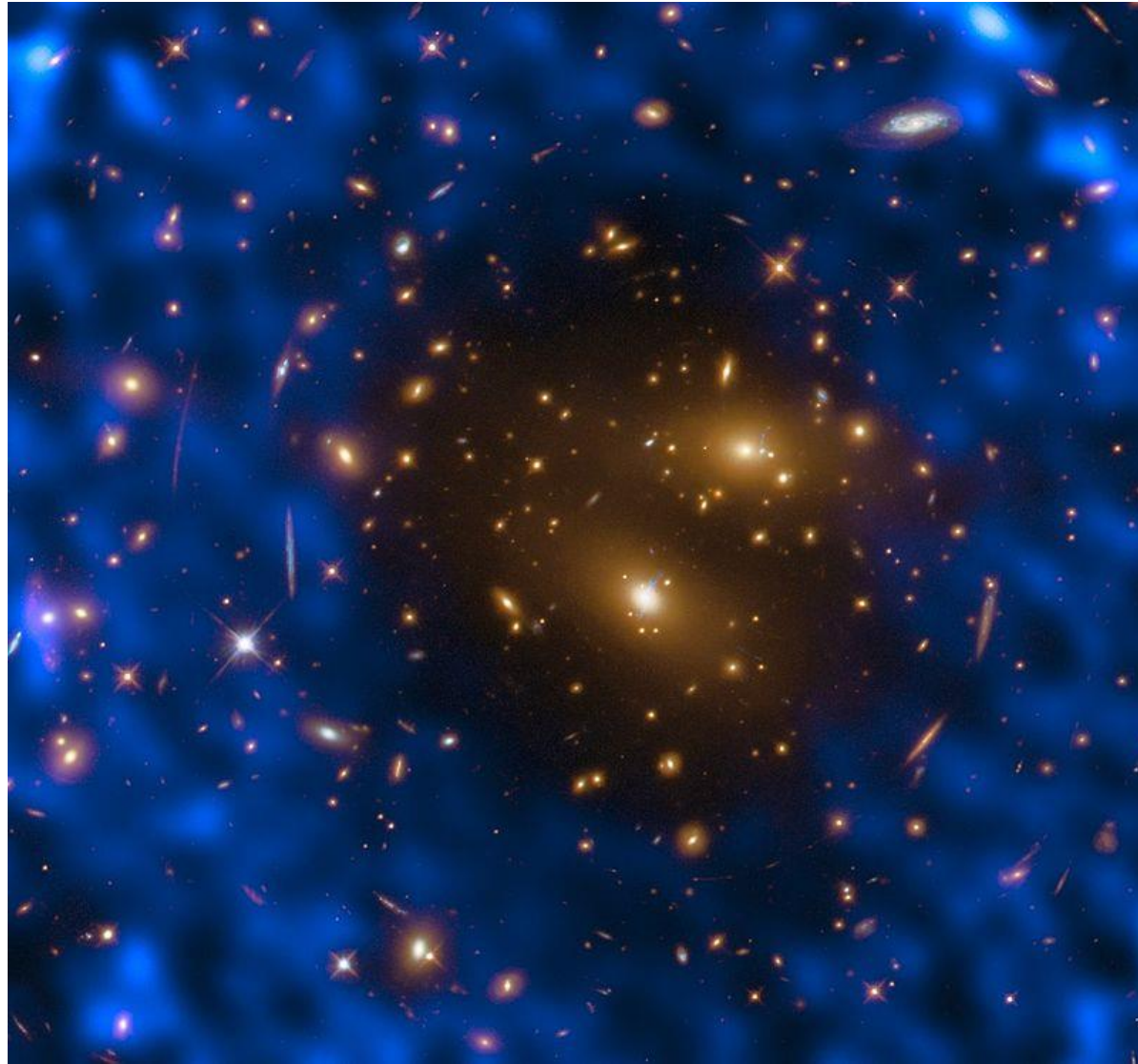
Application: Sunyaev-Zeldovich effect



Compton Scattering



Inverse Compton Scattering



Sunyaev-Zeldovich effect of the most massive galaxy clusters known, [RX J1347.5-1145](#).

V- Conclusions

- When X-rays are scattered off some material, they have different wavelength from the wavelength of the incident X-rays, and this is classically inexplicable.
- Compton used the quantum nature of light to explain the phenomenon.
- Compton's explanation of this effect is important because it demonstrates that light cannot be explained purely as a wave phenomenon. He earned the Nobel price for his discovery.
- The cross section of the light-electron scattering decreases with the increase of the light energy.

Thank you.
