



Doppler broadening free spectroscopy

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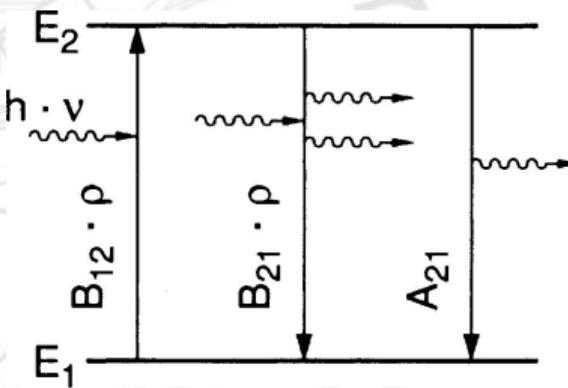
Overview

- ① Introduction to the spectral linewidth
- ② Doppler-free spectroscopy
- ③ Doppler-free spectroscopy at ISOLDE-RILIS
- ④ Conclusions

Overview

- 
- 1 Introduction to the spectral linewidth
 - 2 Doppler-free spectroscopy
 - 3 Doppler-free spectroscopy at ISOLDE-RILIS
 - 4 Conclusions

Absorption, Induced, and Spontaneous Emission



- ▶ probability per second that a molecule absorbs one photon [1]

$$\frac{d}{dt} \mathcal{P}_{12} = B_{12} \rho(\nu) \quad (1)$$

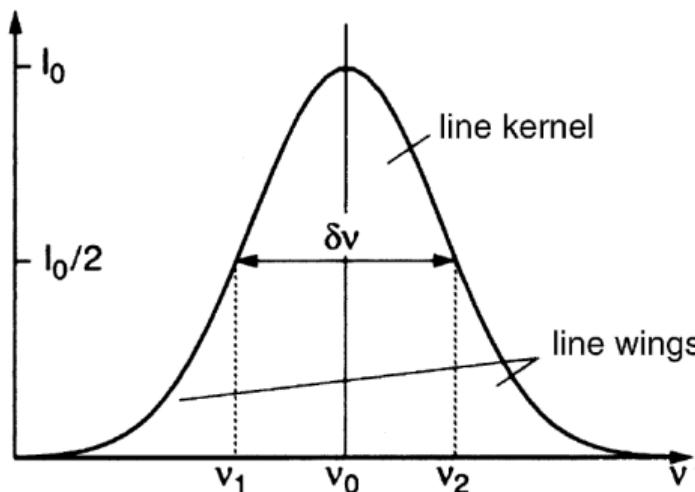
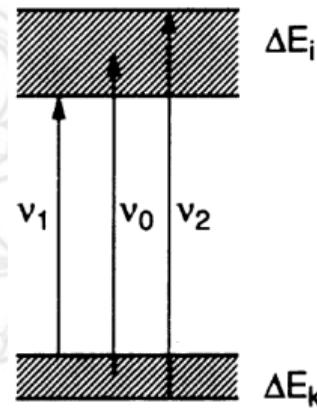
- ▶ probability per second that a molecule emits one induced photon

$$\frac{d}{dt} \mathcal{P}_{21} = B_{21} \rho(\nu) \quad (2)$$

- ▶ probability per second that a molecule emits spontaneously one photon

$$\frac{d}{dt} \mathcal{P}_{21}^{\text{spont}} = A_{21} \quad (3)$$

Line profile

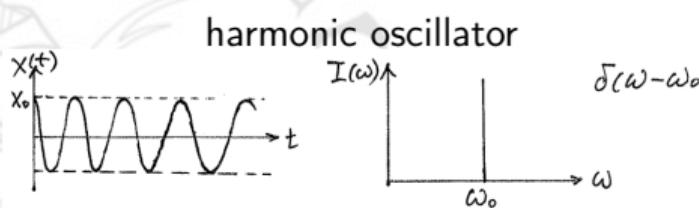


- ▶ The energy levels have a finite width due to the energy–time uncertainty principle
- ▶ The function $I(\nu)$ in the vicinity of ν_0 is called the *line profile*

$$\nu_0 = \frac{E_i - E_k}{h} \quad (4)$$

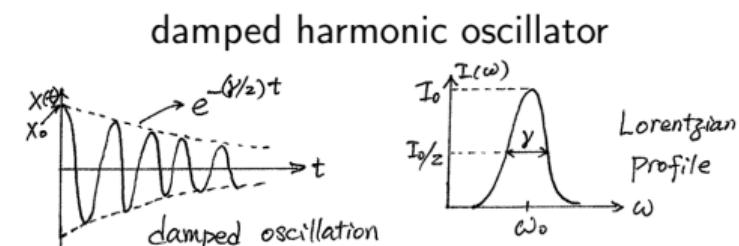
$\delta\nu$ is the FWHM of the line

Natural Linewidth



$$\ddot{x} + \omega_0^2 x = 0 \quad (5)$$

$$x(t) = x_0 \cos \omega_0 t \quad (6)$$



$$\ddot{x} + \gamma \dot{x} + \omega_0^2 x = 0 \quad (7)$$

$$x(t) = x_0 e^{-\frac{\gamma}{2}t} [\cos \omega t + \frac{\gamma}{2\omega} \sin \omega t] \quad (8)$$

$$\omega = \sqrt{\omega_0^2 - \frac{\gamma^2}{4}} \xrightarrow{\gamma \ll \omega_0} \omega_0 \quad (9)$$

$$x(t) = x_0 e^{-\frac{\gamma}{2}t} \cos \omega_0 t \quad (10)$$

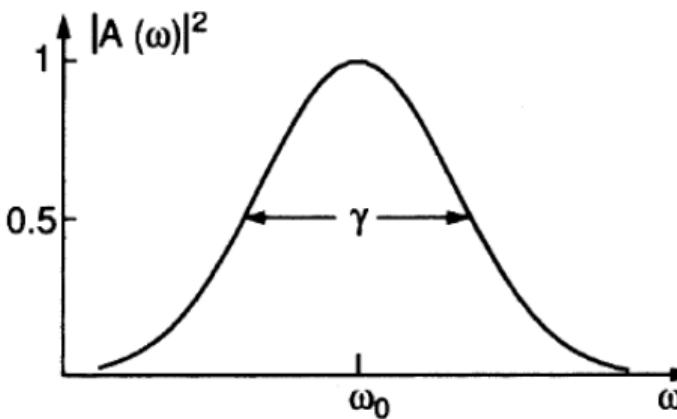
- ▶ atoms modeled as damped harmonic oscillators with central frequency $\omega_0 = 2\pi\nu_0$ that corresponds to $\omega_{ik} = \frac{E_i - E_k}{\hbar}$

Lorentzian line profile

- damped oscillation as superposition of monochromatic oscillations

$$x(t) = \frac{1}{2\sqrt{2\pi}} \int_0^{\infty} A(\omega) e^{i\omega t} d\omega \quad (11)$$

- amplitude $A(\omega)$ as Fourier transform



- real intensity is $I(\omega) \propto A(\omega)A^*(\omega)$

$$I(\omega - \omega_0) = \frac{C}{(\omega - \omega_0)^2 + (\gamma/2)^2} \quad (12)$$

- normalized intensity

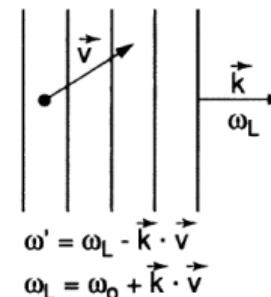
$$L(\omega - \omega_0) = \frac{\gamma/2\pi}{(\omega - \omega_0)^2 + (\gamma/2)^2} \quad (13)$$

with FWHM = γ

Broadening of the natural linewidth

- ▶ Saturation broadening
- ▶ Power broadening
- ▶ Collisional broadening
- ▶ Transit-Time Broadening
- ▶ **Doppler broadening**

Linear Doppler shift



- ▶ In first approximation there is a linear relation between the particle velocity and the frequency shift with respect to an energy level gap $\hbar\omega_0$
- ▶ excited molecule with velocity $\boldsymbol{v} = v_x, v_y, v_z$ emits

$$\omega_e = \omega_0 + \boldsymbol{k} \cdot \boldsymbol{v} \quad (14)$$

- ▶ excitable molecule with velocity $\boldsymbol{v} = v_x, v_y, v_z$ absorbs

$$\omega_a = \omega_0 + \boldsymbol{k} \cdot \boldsymbol{v} \quad (15)$$

Doppler Width

- ▶ At thermal equilibrium, the molecules of a gas follow a Maxwellian velocity distribution
- ▶ Intensity profile for such distribution with central frequency ω_0 is a Gaussian distribution

$$I(\omega) = I_0 \exp \left[- \left(\frac{c(\omega - \omega_0)}{\omega_0 v_p} \right)^2 \right] \quad (16)$$

with most probable velocity $v_p = \sqrt{\frac{2kT}{m}}$

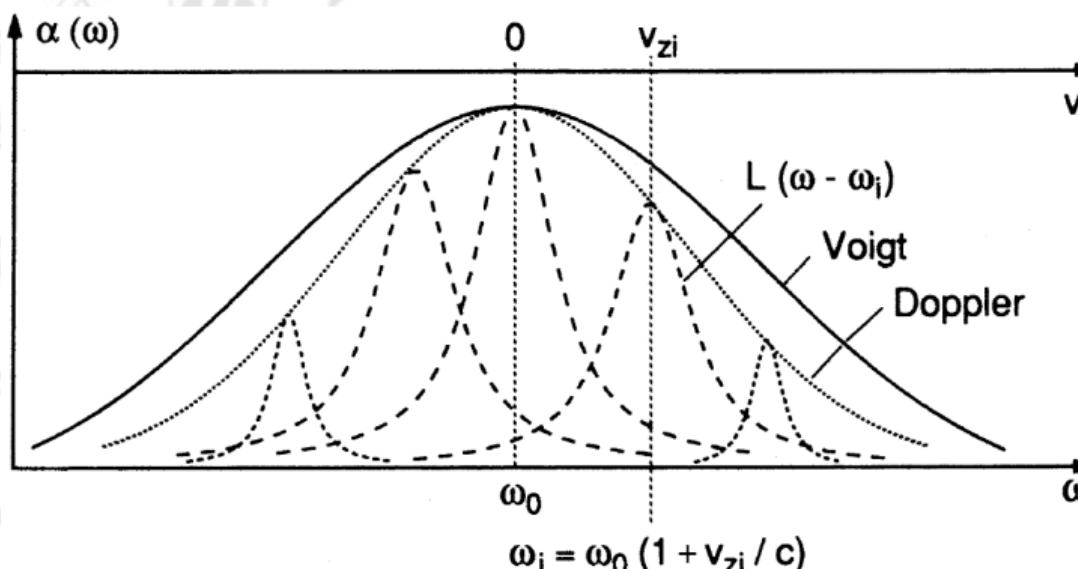
- ▶ Doppler width $\delta\omega_D$ is the FWHM of such Gaussian profile

$$\delta\omega_D = \frac{\omega_0}{c} \sqrt{\frac{8kT \ln 2}{m}} \propto \omega_0 \sqrt{\frac{T}{m}} \quad (17)$$

Voigt Profile

- convolution between Gaussian and Lorentzian

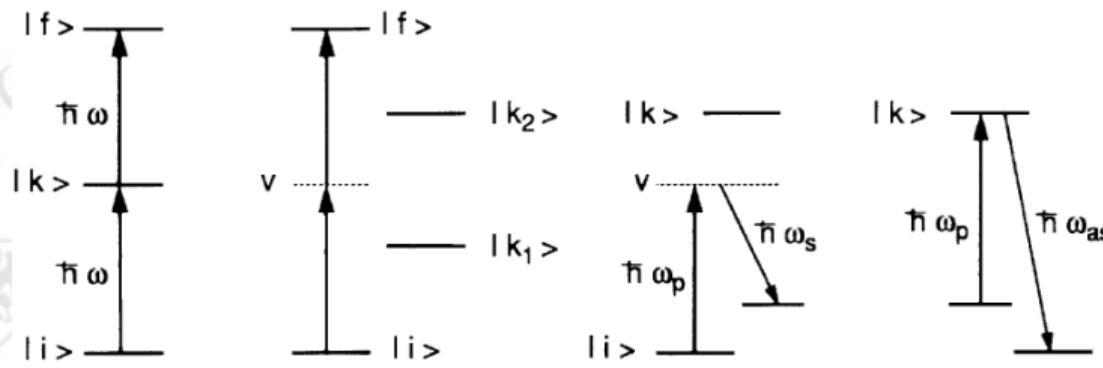
$$I(\omega) = C \int_0^{\infty} \frac{\exp\left(-\left(\frac{c}{v_p} \frac{\omega_0 - \omega'}{\omega_0}\right)^2\right)}{(\omega - \omega')^2 + (\gamma/2)^2} d\omega' \quad \text{with} \quad C = \frac{\gamma N_i c}{2 v_p \pi^{3/2} \omega_0} \quad (18)$$



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Two-photon absorption



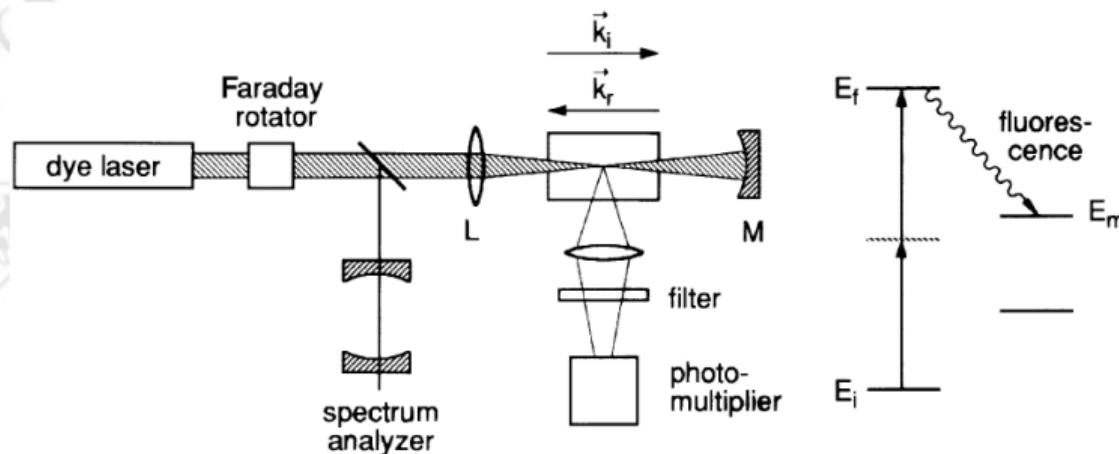
- ▶ Theory from Maria Goeppert-Mayer in 1929 [2]
- ▶ Transition $E_i \rightarrow E_f$ given by several photons of energy ω_i such that $E_f - E_i = \hbar \sum_i \omega_i$

$$A_{if} \propto \frac{\gamma_{if} I_1 I_2}{[\omega_{if} - \omega_1 - \omega_2 - \mathbf{v} \cdot (\mathbf{k}_1 + \mathbf{k}_2)]^2 + (\gamma_{if}/2)^2} \times \text{another factor} \quad (19)$$

- ▶ For $\mathbf{k}_1 = -\mathbf{k}_2$ the Doppler broadening vanishes

$$A_{if} \propto \frac{\gamma_{if} I_1 I_2}{[\omega_{if} - \omega_1 - \omega_2]^2 + (\gamma_{if}/2)^2} \times \text{another factor} \quad (20)$$

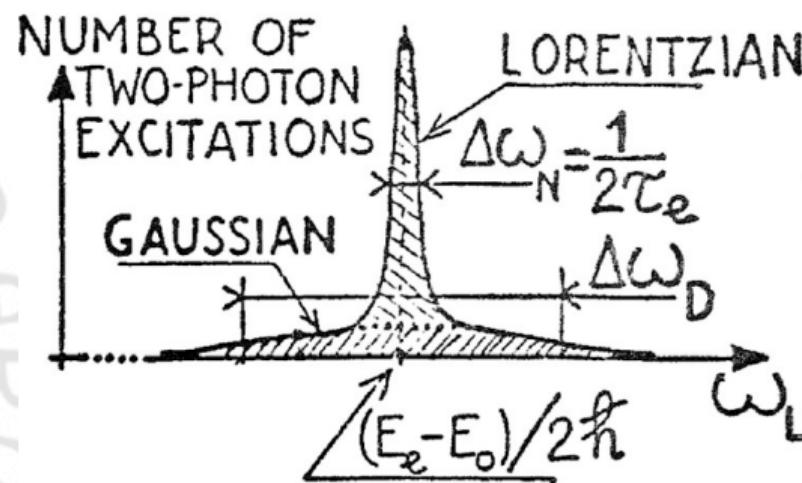
Doppler-free multiphoton spectroscopy



- ▶ Using two photons out of the same laser is disadvantageous

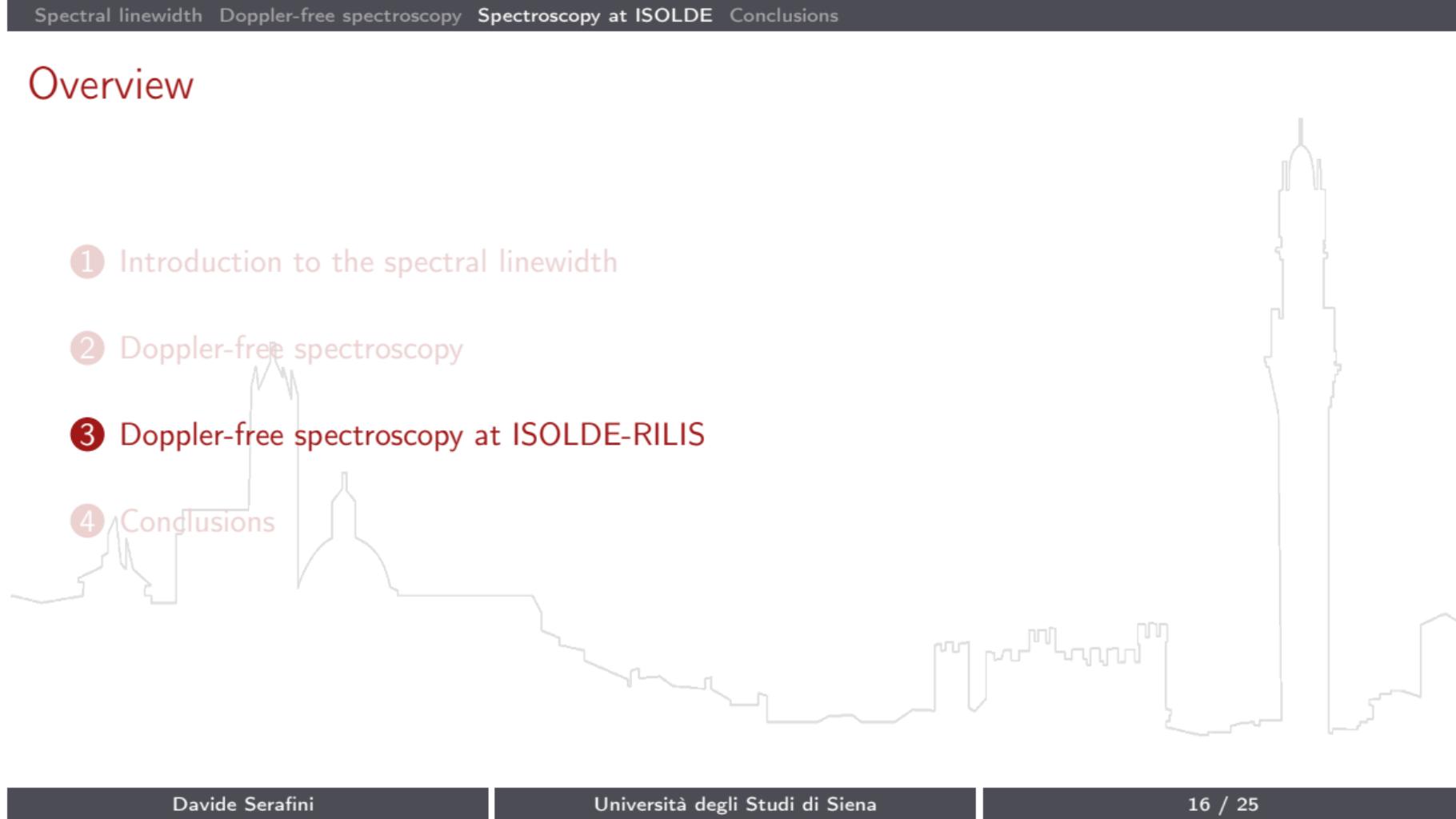
$$2\omega = \frac{E_f - E_i}{\hbar} \quad (21)$$

Two-photon spectrum

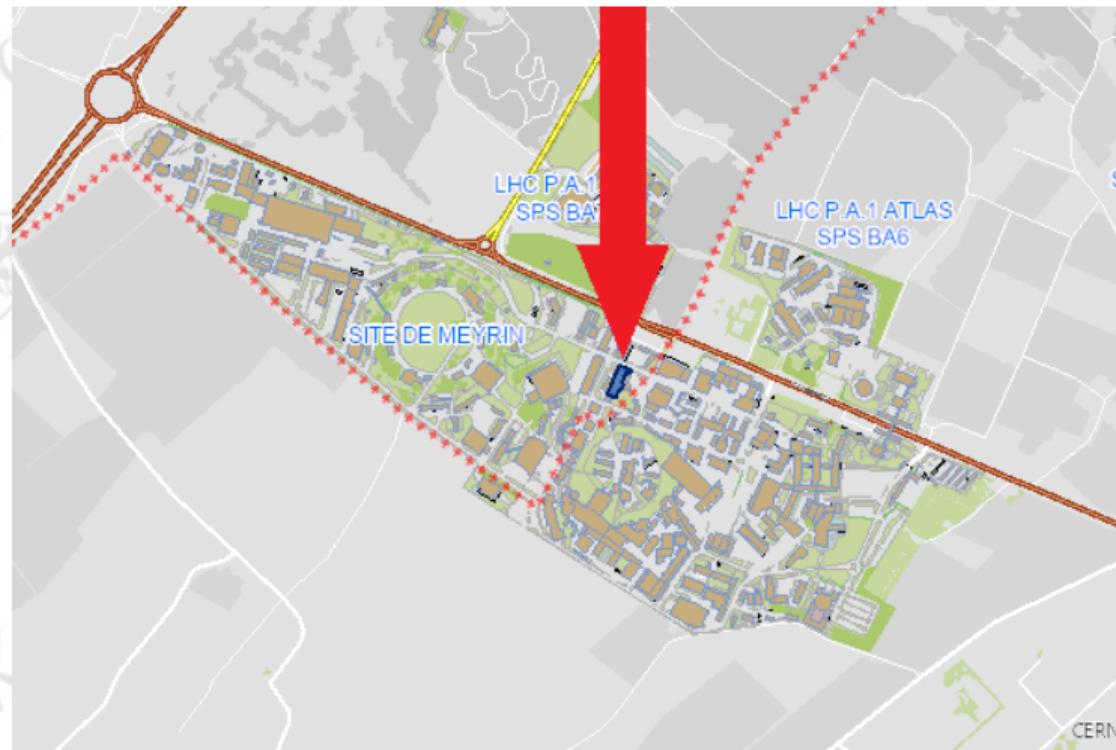


- ▶ The spectrum contains two contributions according to the origin of the two photons:
 - ▶ Doppler-broadened Gaussian ← photons from the same beam
 - ▶ Doppler-free Lorentzian ← photons from counter-propagating beams
- ▶ The Gaussian background can be eliminated using polarized beams

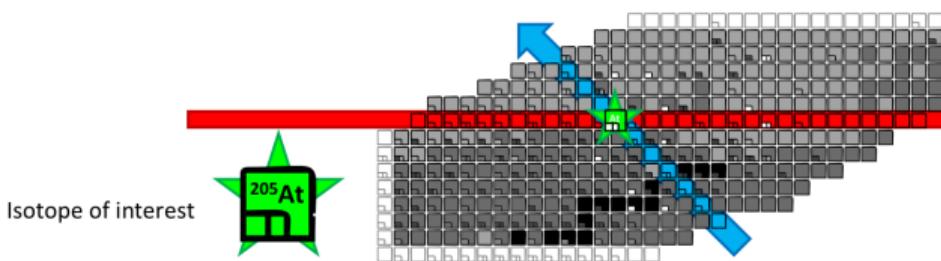
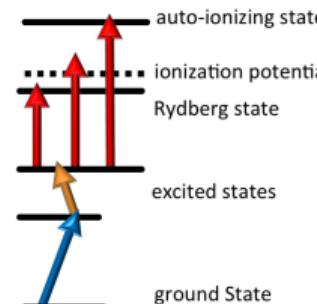
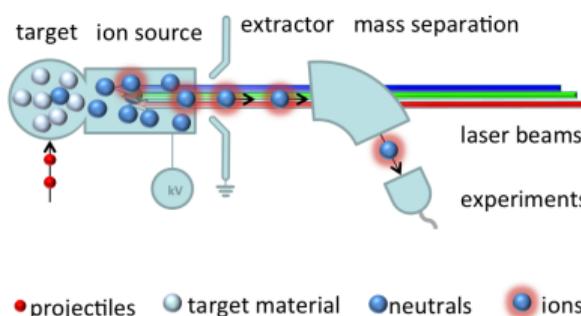
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ISOLDE at CERN

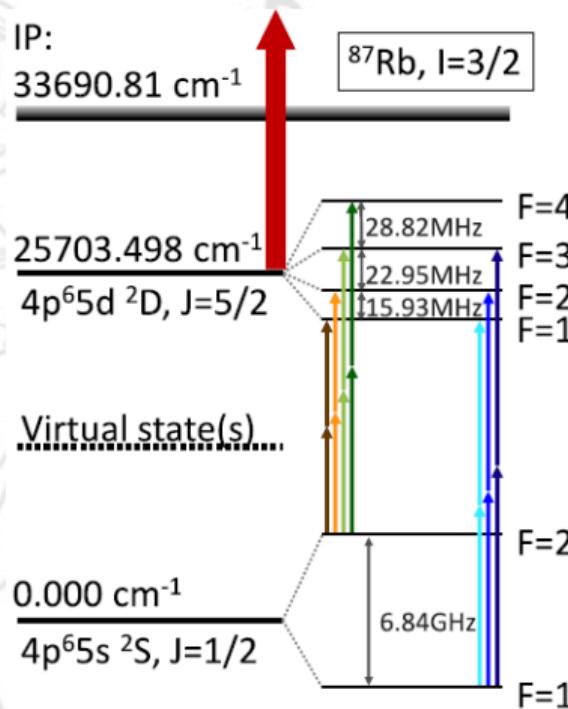


ISOLDE-RILIS



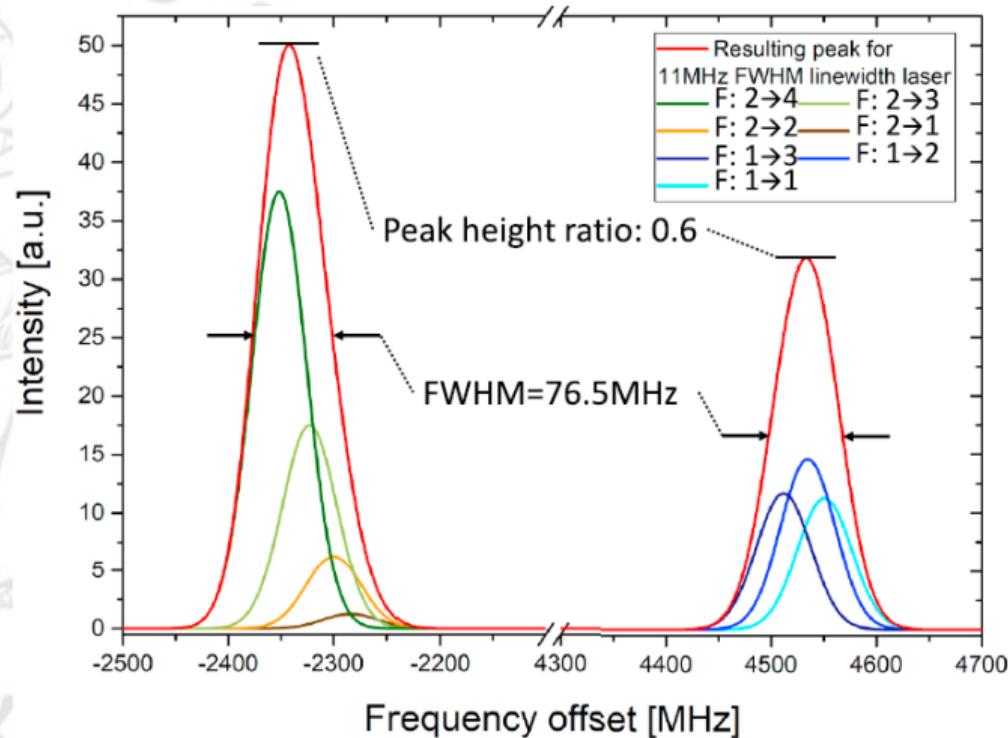
- ▶ Radioactive species are produced in a thick target during its bombardment by proton beam (ISOL technique)
- ▶ Multiple laser beams with specific wavelengths are used to match a series of successive electronic transition energies, unique to that element [3]

Introduction to the experiment



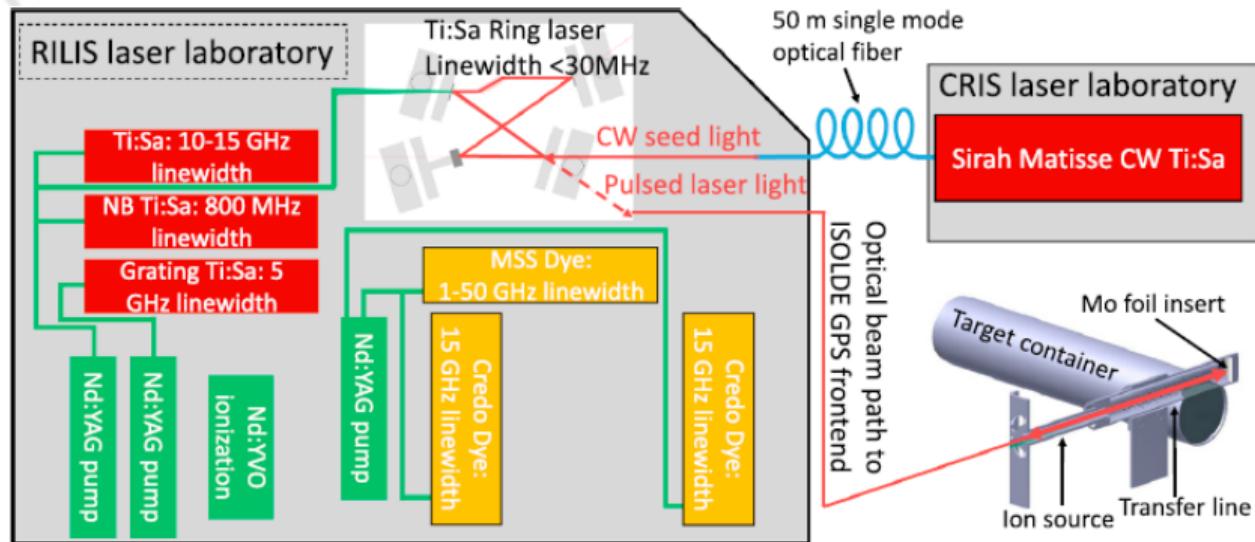
- ▶ Laser ionization scheme for rubidium
- ▶ The hyperfine structure is depicted along with the corresponding shifts (not to scale) and the possible 2-photon transitions
- ▶ The goal of this work is to measure the Doppler-free linewidth, in order to determine the feasibility of laser linewidth limited high-resolution in-source spectroscopy
- ▶ At 2000 K the Doppler broadening is 2.2 GHz for stable rubidium isotopes
- ▶ The frequency of the spectra is given relative to the transition centroid at $25703.498 \text{ cm}^{-1}$

Simulated Spectrum



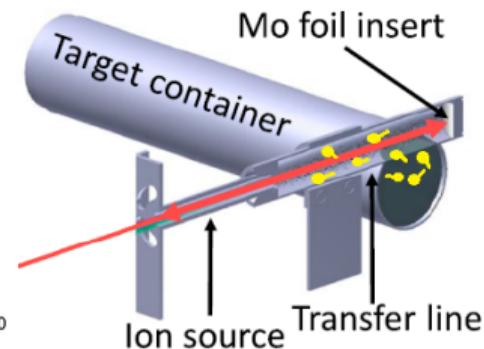
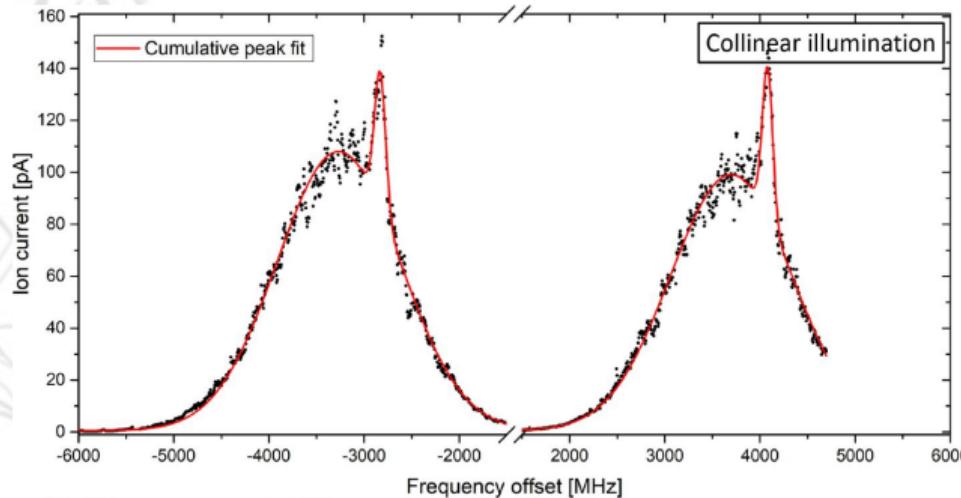
- ▶ The individual transitions can not be resolved under the experimental conditions
- ▶ The 6.84 GHz split of the ground level can be resolved
- ▶ Assumed laser linewidth ~ 10 MHz

Collinear Doppler-free 2-photon resonance ionization



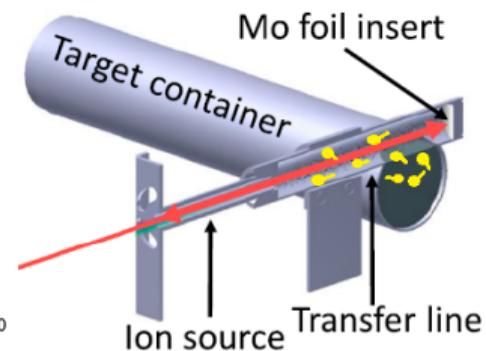
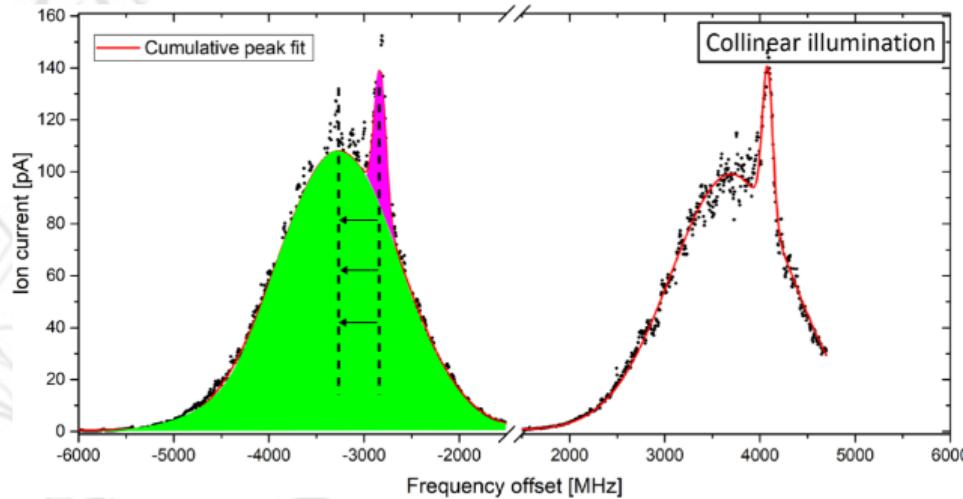
- ▶ First demonstration of Doppler-free 2-photon in-source laser spectroscopy at the ISOLDE-RILIS [4]
- ▶ Mirror realized with a flat molybdenum foil

ISOLDE two-photons results



- ▶ narrow Doppler-free signal
- ▶ Doppler broadened and shifted background:
 - ▶ the atom source, transfer line and ion source geometry results in an atomic sample with predominantly velocity components towards the extraction electrode
 - ▶ the reflectivity is less than 100 %

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Conclusions

- ▶ Doppler broadening is a major limitation in laser spectroscopy
- ▶ Doppler-free spectroscopy can be achieved through two-photon absorption
- ▶ At CERN-ISOLDE this technique was used to measure the spectral profiles of rubidium atoms

A faint watermark of the Siena University logo is visible on the left side of the slide. It features a stylized figure of a saint or angel standing next to a shield with a heraldic emblem. The text "SAPIENS" is curved along the top edge, and "CXXV" is at the bottom.

Thank you for the attention

References I

- [1] Wolfgang Demtröder. "Absorption and Emission of Light". In: *Laser Spectroscopy 1: Basic Principles*. Berlin, Heidelberg: Springer Berlin Heidelberg, 2014, pp. 5–74. ISBN: 978-3-642-53859-9. DOI: 10.1007/978-3-642-53859-9_2. URL: https://doi.org/10.1007/978-3-642-53859-9_2.
- [2] Wolfgang Demtröder. *Laser Spectroscopy 2: Experimental Techniques*. Springer Berlin Heidelberg, 2015. Chap. 2, pp. 118–131. ISBN: 9783662446416. DOI: 10.1007/978-3-662-44641-6. URL: <http://dx.doi.org/10.1007/978-3-662-44641-6>.
- [3] URL: <https://rilia-web.web.cern.ch/> (visited on 11/25/2024).

References II

- [4] K. Chrysalidis et al. "First demonstration of Doppler-free 2-photon in-source laser spectroscopy at the ISOLDE-RILIS". In: *Nuclear Instruments and Methods in Physics Research Section B: Beam Interactions with Materials and Atoms* 463 (2020), pp. 476–481. ISSN: 0168-583X. DOI: doi.org/10.1016/j.nimb.2019.04.020. URL: <https://www.sciencedirect.com/science/article/pii/S0168583X19302046>
- [5] Peter W. Milonni and Joseph H. Eberly. "Laser Oscillation: Gain and Threshold". In: *Laser Physics*. John Wiley & Sons, Ltd, 2010. Chap. 4, pp. 141–173. ISBN: 9780470409718. DOI: <https://doi.org/10.1002/9780470409718.ch4>. eprint: <https://onlinelibrary.wiley.com/doi/pdf/10.1002/9780470409718.ch4>. URL: <https://onlinelibrary.wiley.com/doi/abs/10.1002/9780470409718.ch4>.

References III

- [6] Peter W. Milonni and Joseph H. Eberly. "Laser Oscillation: Power and Frequency". In: *Laser Physics*. John Wiley & Sons, Ltd, 2010. Chap. 5, pp. 175–228. ISBN: 9780470409718. DOI: <https://doi.org/10.1002/9780470409718.ch5>. eprint: <https://onlinelibrary.wiley.com/doi/pdf/10.1002/9780470409718.ch5>. URL: <https://onlinelibrary.wiley.com/doi/abs/10.1002/9780470409718.ch5>.

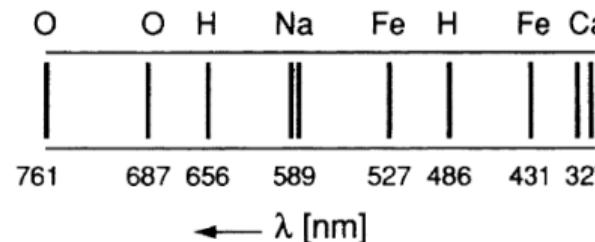


Back-up

Questions

- ▶ How long is a laser pulse?
- ▶ Is it possible to have no crossing of the forward and back propagating beams in the source
- ▶ What is the maximum time interval to have two-photon absorption and not one photon absorbed after the other?

Spectroscopy



Fraunhofer lines in the spectrum of the sun

Spectroscopy

- ▶ Spectroscopic investigations allow us to study atoms and molecules structures [1]

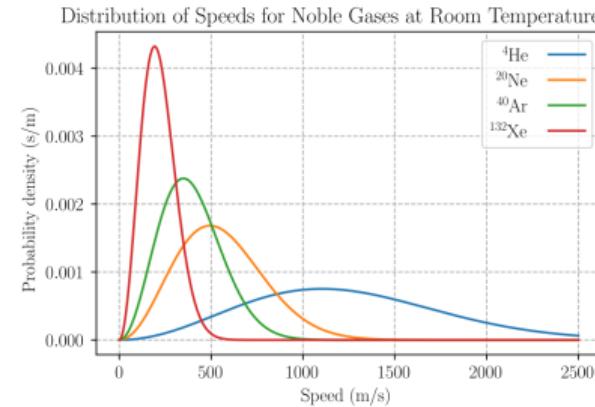
Maxwell-Boltzmann distribution

- ▶ this implies that the momenta p and kinetic energies E_{kin} of the various particle species follow a Maxwell-Boltzmann distribution

Maxwell-Boltzmann equations

$$n(p)dp = \sqrt{2\pi}N \left(\frac{1}{\pi m K_B T} \right)^{3/2} p^2 e^{-\frac{p^2}{2mK_B T}} dp \quad (22)$$

where N denotes the total number of particles in the system and m their mass



Cavity Modes

assumptions [1]

- ▶ cubic cavity with the sides L at the temperature T
- ▶ walls of the cavity absorb and emit electromagnetic radiation
- ▶ at thermal equilibrium $\Rightarrow P_a(\omega) = P_e(\omega) \quad \forall \omega$

stationary radiation field E

$$\mathbf{E} = \sum_p \mathbf{A}_p \exp[i(\omega_p t - \mathbf{k}_p \cdot \mathbf{r})] + \text{c.c.} \quad (23)$$

Thermal Radiation and Planck's Law

Rayleigh-Jeans law

- ▶ classical oscillators applied to the cubic cavity

$$\rho(\nu)d\nu = n(\nu)kTd\nu = \frac{8\pi\nu^2k}{c^3} Td\nu \quad (24)$$

Planck's radiation law

- ▶ only discrete amounts $qh\nu$ with q integer can be absorbed/emitted by the radiation field

$$\rho(\nu)d\nu = \frac{8\pi\nu^2}{c^3} \frac{h\nu}{e^{h\nu/kT} - 1} d\nu \quad (25)$$

Spectroscopy Instrumentation

- ▶ the nuclear time scale is

$$\tau_n = \frac{E_n}{L} F \quad (26)$$

where E_n is the nuclear energy reservoir, L is the luminosity and F the fraction of reservoir available

Power Broadening

- ▶ Fermi-Dirac distribution must be used

$$n(p)dp = \frac{8\pi p^2}{h^3} \left(\frac{1}{1 + e^{-\eta + E_{\text{kin}}/K_B T}} \right) dp \quad (27)$$

- ▶ in the case of electron degeneracy the electron pressure does not depend on the temperature

$$P = 1.0036 \times 10^{13} \left(\frac{\rho}{\mu_e} \right)^{5/3} \quad (28)$$

Other sources of spectral line broadening

- ▶ collisional broadening
- ▶ power broadening

Quadratic Doppler shift

The momentum \mathbf{p}_i of an atom absorbing a photon with momentum $\hbar\mathbf{k}$ and energy $\hbar\omega_{ik} \simeq E_k - E_i$ changes to

$$\mathbf{p}_k = \mathbf{p}_i + \hbar\mathbf{k} \quad (29)$$

Using relativistic energy considerations:

$$\hbar\omega_{ik} = \sqrt{p_k^2 + (M_0 c^2 + E_k)^2} - \sqrt{p_i^2 c^2 + (M_0 c^2 + E_i)^2} \quad (30)$$

Use Taylor expansion:

$$\omega_{ik} = \underbrace{\omega_0}_{\text{at rest}} + \underbrace{\mathbf{k} \cdot \mathbf{v}_i}_{\text{linear Doppler}} - \underbrace{\frac{-\omega_0 v_i^2}{2c^2} + \frac{\hbar\omega_0^2}{2Mc^2}}_{\text{quadratic Doppler} \rightarrow \text{recoil}} + \dots \quad (31)$$

cross section

$$\text{stimulated emission rate} = \sigma(\nu) \frac{\text{number of incident photons}}{\text{area} \cdot \text{time}} \quad (32)$$

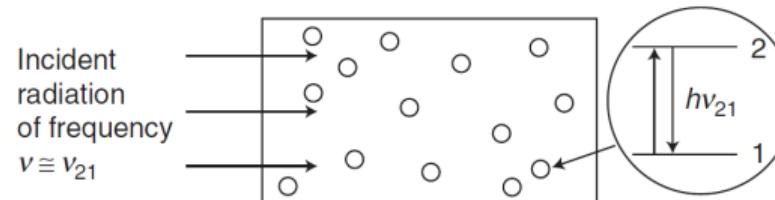
with

$$\sigma(\nu) = \frac{\lambda^2 A_{21}}{8\pi} S(\nu) \quad (33)$$

it depends on:

- ▶ spontaneous emission rate A_{21}
- ▶ lineshape function $S(\nu)$
- ▶ transition wavelength λ

Lasing



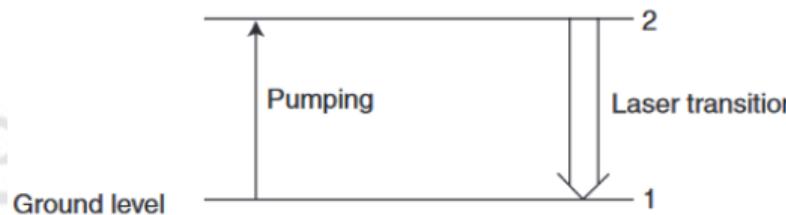
Suppose a beam of photons of energy $h\nu_{21}$ passing through a medium:

- ▶ $N_2 < N_1 \rightarrow$ absorption $\rightarrow g < 0$
- ▶ $N_2 > N_1 \rightarrow$ amplification $\rightarrow g > 0$

Having defined the gain coefficient g as

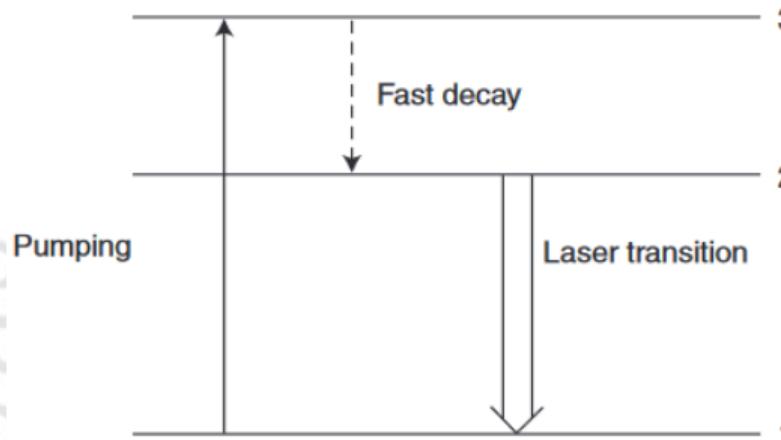
$$g(\nu) = \sigma(\nu) \left(N_2 - \frac{g_2}{g_1} N_1 \right) I_\nu \quad (34)$$

Two-level scheme



- ▶ Just two levels do not allow for population inversion
- ▶ The pump will also contribute to the depletion of the higher level

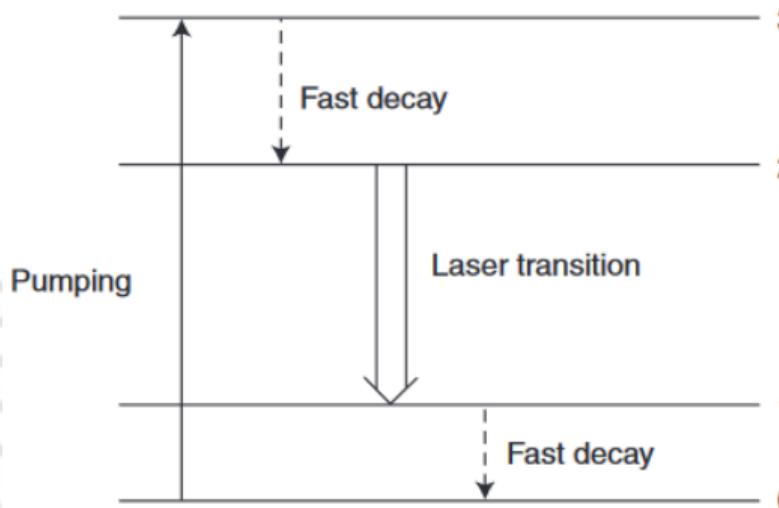
Three-level scheme



- ▶ Population inversion is possible having decoupled the pump from the laser transition

$$\bar{N}_2 - \bar{N}_1 = \frac{P - \Gamma_{21}}{P + \Gamma_{21}} N \quad (35)$$

Four-level scheme



- ▶ Population inversion is guaranteed ($\bar{N}_2 - \bar{N}_1 > 0$ always) choosing a medium with spontaneous emission coefficients such that $\Gamma_{10} > \Gamma_{21}$

$$\bar{N}_2 - \bar{N}_1 = \frac{\Gamma_{10} - \Gamma_{21}}{\Gamma_{10}\Gamma_{21} + P\Gamma_{21} + P\Gamma_{10}} PN \xrightarrow{\Gamma_{10} \gg \Gamma_{21}, P} \frac{P}{P + \Gamma_{21}} N \quad (36)$$

Equations ruling the states' populations

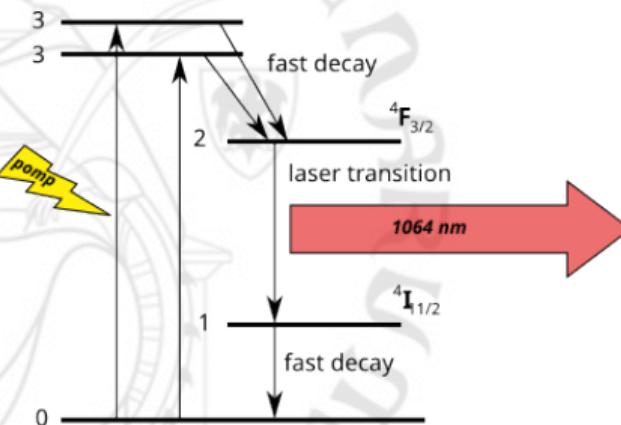
$$\frac{dN_0}{dt} = -PN_0 + \Gamma_{10}N_1 \quad (37)$$

$$\frac{dN_1}{dt} = -\Gamma_{10}N_1 + \Gamma_{12}N_2 + \sigma(\nu)(N_2 - N_1)\Phi_\nu \quad (38)$$

$$\frac{dN_2}{dt} = PN_0 - \Gamma_{21}N_2 - \sigma(\nu)(N_2 - N_1)\Phi_\nu \quad (39)$$

assuming that the decay $3 \rightarrow 2$ is much faster than the others so that $N_3 \approx 0$

Nd:YAG



- ▶ Neodymium-doped Yttrium Aluminum Garnet → $\text{Nd}:\text{Y}_3\text{Al}_5\text{O}_{12}$
- ▶ Neodymium impurities are in a +3 oxidation state (Nd(III))
- ▶ Nd(III) typically replace about 1 % of the Y atoms in the crystal structure of the YAG
- ▶ Nd:YAG laser is approximately a four-level system

Saturation

- ▶ Assume a two-level system with no pump
- ▶ The steady state solutions for the number of atoms in the two energy levels are

$$\bar{N}_2 = \frac{\frac{1}{2}I_\nu/I_\nu^{\text{sat}}}{1 + I_\nu/I_\nu^{\text{sat}}} N \quad (40)$$

$$\bar{N}_1 = \frac{1 + \frac{1}{2}I_\nu/I_\nu^{\text{sat}}}{1 + I_\nu/I_\nu^{\text{sat}}} N \quad (41)$$

- ▶ I_ν is the intensity of the photons with energy $h\nu$ in the medium
- ▶ I_ν is not related to the pump intensity, that is not considered here
- ▶ the saturation intensity

$$I_\nu^{\text{sat}} = \frac{h\nu A_{21}}{2\sigma(\nu)} \quad (42)$$

provides the measure of whether a given field intensity I_ν is large or small in terms of its ability to saturate the transition $1 \rightarrow 2$

Saturation

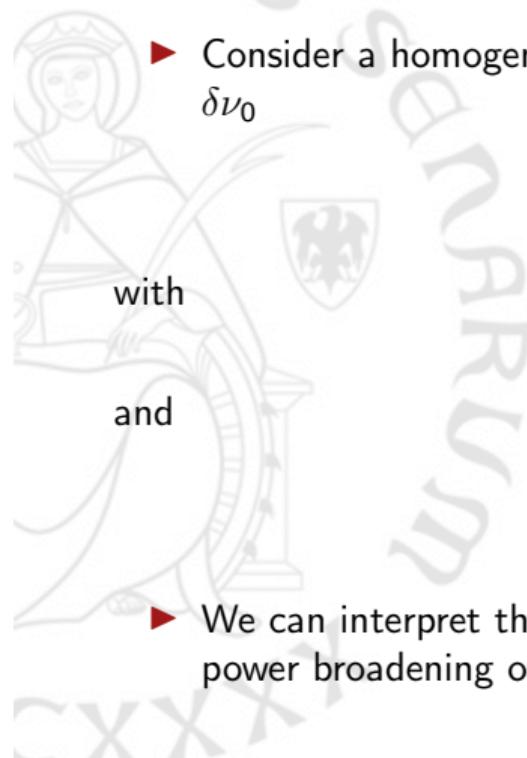
The absorption coefficient is

$$a(\nu) = \sigma(\nu)(\bar{N}_1 - \bar{N}_2) = \frac{\sigma(\nu)N}{1 + I_\nu/I_\nu^{\text{sat}}} \equiv \frac{a_0(\nu)}{1 + I_\nu/I_\nu^{\text{sat}}} \quad (43)$$

- ▶ $I_\nu \ll I_\nu^{\text{sat}} \rightarrow \bar{N}_1 \approx N, \quad \bar{N}_2 \approx 0, \quad a(\nu) \approx a_0(\nu)$
- ▶ $I_\nu \gg I_\nu^{\text{sat}} \rightarrow \bar{N}_1 \approx N/2, \quad \bar{N}_2 \approx N/2, \quad a(\nu) \approx 0$
- ▶ as I_ν grows the stimulated absorption and emission processes overshadow the spontaneous emission ones

Power broadening

- ▶ Consider a homogeneously broadened transition having a Lorentzian lineshape of width $\delta\nu_0$


$$a(\nu) = \frac{a_0(\nu_0)\delta\nu_0^2}{(\nu - \nu_0)^2 + \delta\nu_0'^2} \quad (44)$$

with

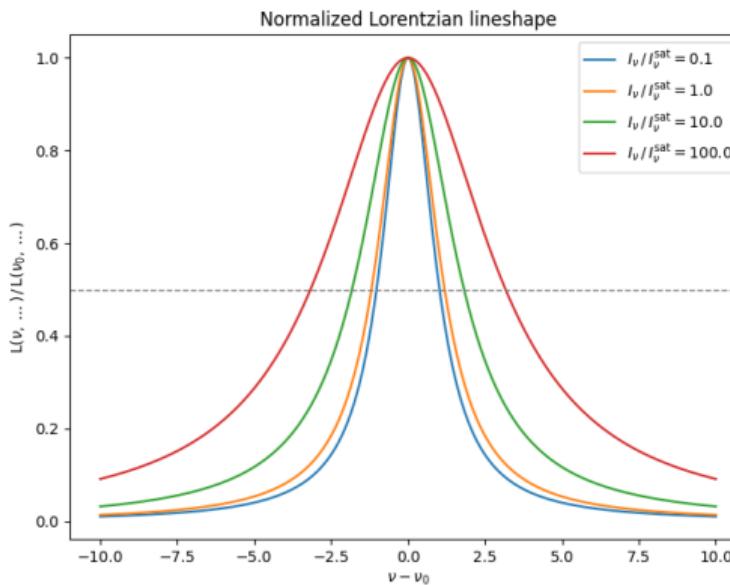
and

$$\delta\nu_0' = \delta\nu_0 \sqrt{1 + I_\nu/I_\nu^{\text{sat}}} \quad (45)$$

$$a_0(\nu_0) = \frac{\lambda^2 A_{21} N}{8\pi^2 \delta\nu_0} \quad (46)$$

- ▶ We can interpret the saturation of the transition with increasing intensity as an effective power broadening of the linewidth

Power broadening



- ▶ In general the saturation intensity I_ν^{sat} will depend on both upper- and lower-level decay rates associated with collisional as well as radiative processes, and it can also depend on the level degeneracies
- ▶ what matters is that absorption $a(\nu)$ and gain $g(\nu)$ coefficients saturates increasing the intensity as $\sqrt{1 + I_\nu / I_\nu^{\text{sat}}}$, whatever the form of I_ν^{sat}

Saturation for a four-level scheme

- ▶ We consider a four-level scheme with pumping rate P
- ▶ Gain coefficient

$$g(\nu) = \frac{g_0(\nu)}{1 + I_\nu/I_\nu^{\text{sat}}} \quad (47)$$

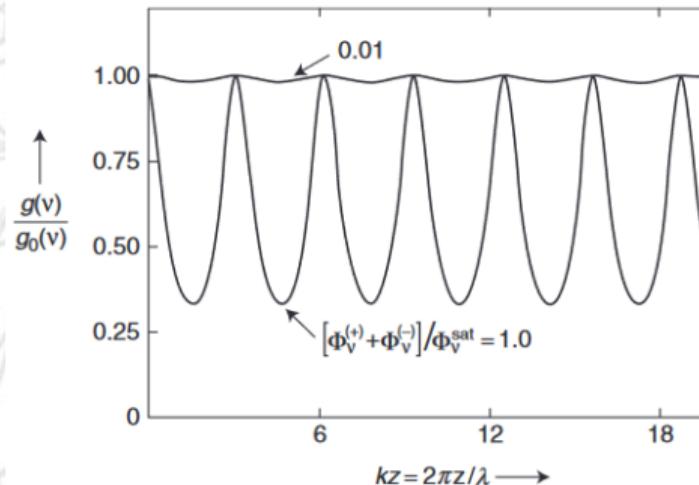
- ▶ saturation intensity
- ▶ saturation flux

$$I_\nu^{\text{sat}} = h\nu\Phi_\nu^{\text{sat}} \quad (48)$$

$$\Phi_\nu^{\text{sat}} = \frac{P + \Gamma_{21}}{\sigma(\nu)} \quad (49)$$

the saturation coefficients depends also on the pumping rate P

Spatial Hole Burning

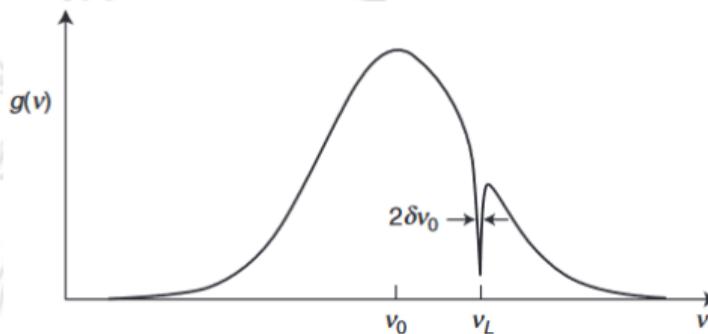


- In most lasers there is a standing wave rather than a traveling one

$$\Phi_\nu = \Phi_\nu^{(+)} + \Phi_\nu^{(-)} \quad (50)$$

- Therefore, the intensity of the field I_ν is affected by the interference of the two oppositely propagating waves
- The gain $g(\nu)$ has a term $\sin^2 kz$ that modulates its maximum, small signal value $g_0(\nu)$ along the cavity \rightarrow spatial hole burning

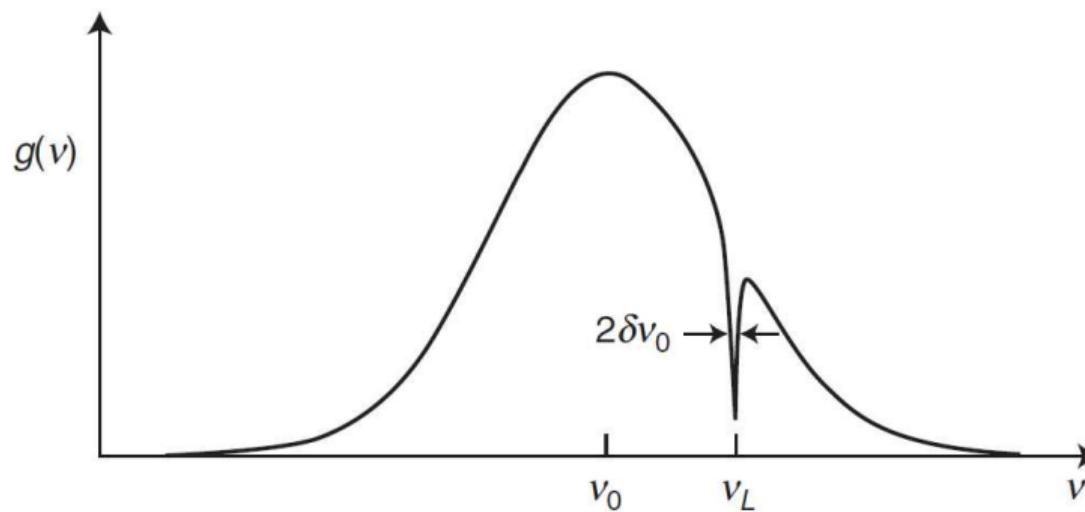
Spectral Hole Burning



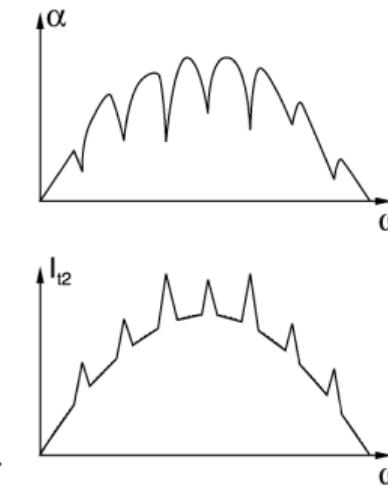
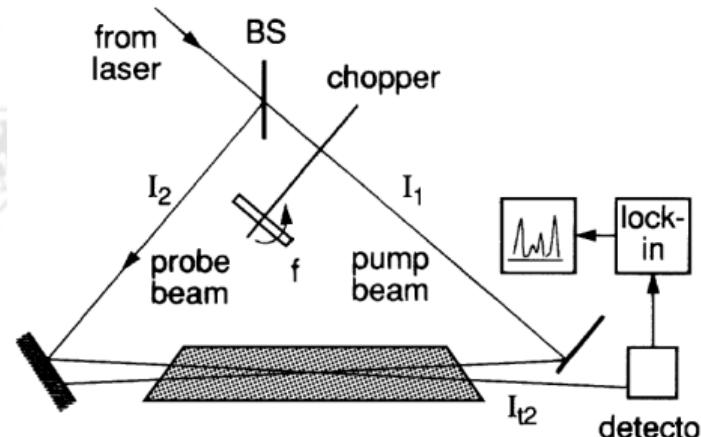
- ▶ Consider an inhomogeneously broadened medium where the atoms have different central transition frequencies ν'_0
- ▶ The cavity mode frequency is ν
- ▶ The absorption or gain is more strongly saturated for spectral packets with frequency $\nu'_0 \approx \nu_0$
- ▶ The selective saturation leads to spectral hole burning in the gain curve

Hole burning

- ▶ Spectral hole burning is the origin of the Lamb dip in Doppler-broadened lasers [5]
- ▶ This dip in output power at line center was predicted by W. E. Lamb, Jr. in 1963, and is called the Lamb dip [6]

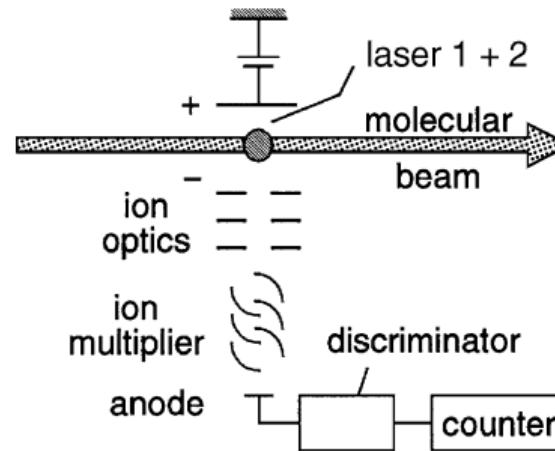


Saturation spectroscopy



► <https://www.youtube.com/watch?v=hY21ndCeZ9I>

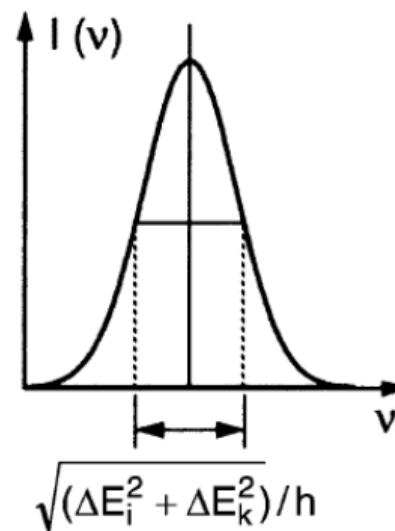
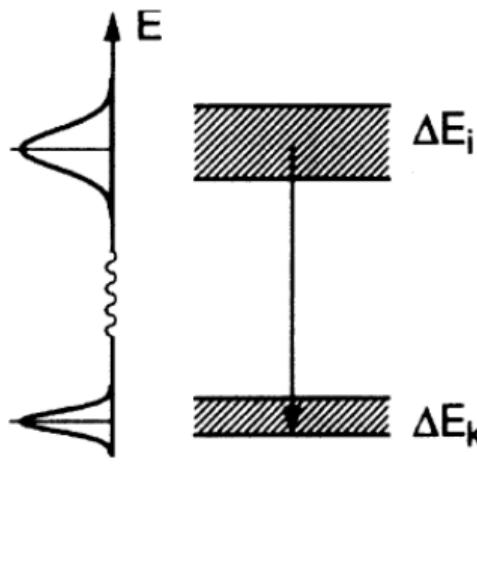
Experimental arrangement for ionization spectroscopy with pulsed laser



- ▶ experimental arrangement for photoionization spectroscopy in a molecular beam
- ▶ ionization is the most sensitive detection method
- ▶ suitable for small absorption coefficients α

$$I = I_0 e^{-\alpha(\omega)z} \quad (51)$$

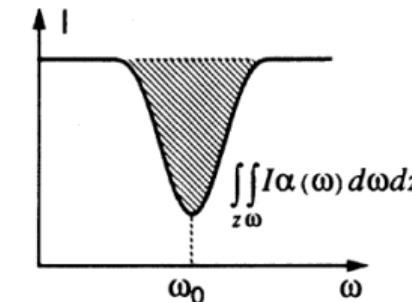
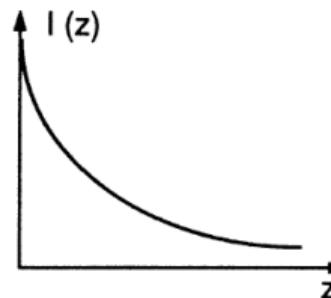
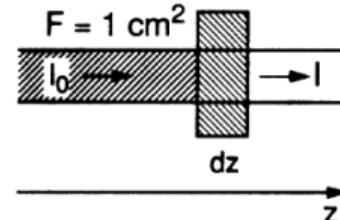
Linewidth and Lifetime



$$\omega_{ik} = \frac{E_i - E_k}{\hbar} \quad (52)$$

$$\delta\omega = \frac{\sqrt{\Delta E_i^2 + \Delta E_k^2}}{\hbar}$$

Natural linewidth of absorbing transitions



$$dI = -\alpha I dz \quad (53)$$

$$\alpha_{ik}(\omega) = \sigma_{ik}(\omega) \left[N_i - \frac{g_i}{g_k} N_k \right] \quad (54)$$

$$\alpha(\omega) = \frac{Ne^2}{4\epsilon_0 mc} \frac{\gamma}{(\omega - \omega_0)^2 + (\gamma/2)^2} \quad (55)$$

Lorentzian profile with FWHM $\Delta\omega_n = \gamma$